# Fragmentation of the low energy octupole phonon in (196)Pt and ... 

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## THE FLORIDA STATE UNIVERSITY <br> COLLEGE OF ARTS AND SCIENCES

# FRAGMENTATION OF THE LOW ENERGY OCTUPOLE PHONON IN ${ }^{196} \mathbf{P t}$ AND INVERSE KINEMATICS PROTON SCATTERING FROM THE RADIOACTIVE BEAM ${ }^{20} \mathrm{O}$ 

By
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A Dissertation submitted to the
Department of Physics
in partial fulfillment of the requirements for the degree of Doctor of Philosophy

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The members of the Committee approve the Dissertation of J. Keith Jewell defended on September 8, 1997.


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This work is dedicated to my wife Barb, who has inspired me beyond my own expectations, and has taught me how to live fully.

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#### Abstract

The results from two separate proton scattering experiments are presented. First, we present positive identification of a $J^{\pi}=3^{-}$state at $2423.4(4) \mathrm{keV}$ in ${ }^{196} \mathrm{Pt}$ from a ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ reaction at 12.7 MeV performed at Florida State University. The identification, based on known proton and $\gamma$-ray angular distributions, indicates that the low energy octupole state is fragmented between at least two states. Octupole fragmentation in ${ }^{196} \mathrm{P}$ t is discussed in terms of the $\mathrm{O}(6)$ symmetry of the IBM. Second, we present angular distributions of protons scattered from the ground state and $2_{1}^{+}$state of the singly closed shell uncleus ${ }^{20} \mathrm{O}$, measured in inverse kinematics at $30 \mathrm{MeV} / \mathrm{A}$ at the National Superconducting Cyclotron Laboratory. We extract the effective RMS quadrupole deformation parameter $\beta_{2}$ for proton scattering by comparison with a coupled channels calculation. The known electromagnetic $B\left(E 2 ; 0_{g s}^{+} \rightarrow \mathscr{2}_{1}^{+}\right)$is used in conjunction with our deduction of $\beta_{2}$ from proton scattering to extract a ratio of neutron to proton multipole matrix elements $M_{n} / M_{p}=2.8 \pm 0.40$ for the $0_{g s}^{+} \rightarrow \vartheta_{1}^{+}$transition. The present value of $M_{n} / M_{p}$ indicates a large neutron contribution to the $\geq_{1}^{+}$state excitation due to the closed $Z=8$ proton shell.


## CHAPTER 1

## INTRODUCTION

The low energy behavior of nuclei can be categorized, in the most basic way, as either single-particle-like or collective. The single-particle description is strongly related to a shell model formalism reminiscent of descriptions applied to electron behavior in atomic systems. The emphasis of the shell model is on discrete orbits or angular momentum states successively filled by individual nucleons. The simple collective description, in contrast, relates a many-body uuclear system to a classical liquid drop shape without specifying the behavior of single nucleons. Collective models are most useful in describing the overall shape of the nucleus, and time dependent deviatious from that equilibrium shape. There are not, however, clear and sharp distinctions between collective and single-particle behavior. It is now understood that most nuclei exhibit both types of behavior to varying degrees. Purely collective or single-particle descriptions represent only the extremes of a continuous variety of uuclear behavior.

The shell model description was first to gain wide acceptance in the scientific community for a number of reasous. Existing shell model descriptions of electrous in atomic systems made a parallel development for nuclear systems quite attractive. Early experimental evidence also indicated that a shell model gave a valid description of many nuclear properties. The existence of strongly peaked isotopic abundances of nuclei clustered about specific nucleon numbers and large jumps seeu in nucleon separation energies pointed to an underlying shell structure for nucleous. This experimental evidence hinted at the existence of "magic" numbers, which are indicative of completely filled nuclear shells. By introducing
a strong spin-orbit coupling to a simple harmonic oscillator potential, we obtain the shell structure used today, having magic numbers at $2,8,20,40,50,82$, and 126 [Fee49, Hax49, May49, May50, Jeu52, May55].

Measurements of neutron capture cross sections and quadrupole moments lent further confidence in the shell model description [May48]. Early forms of the shell model were successful in explaining spins and parities of ground states in many nuclei within a few nucleons of the magic umbers [Hax49, May.55]. Lastly, the shell model gained acceptance because early experimental studies focused on light nuclei, for which the shell model is more easily applied. Since many studies of nuclei often require probing the nucleus above the Coulomb barrier, advanced accelerator technology was needed to deliver beams of sufficient energy to study some of the heavier nuclei.

As more experimental evidence became available, however, the shell model could not reproduce the properties of nuclei away from the magic numbers. Nuclei residing far away from closed shells exhibited behavior which could not be easily explained by a nuclear shell structure. The most striking evidence that the shell model was an incomplete picture came from study of electric quadrupole moments in nuclei away from closed shells. The electric quadrupole moment is a measure of the quadrupole deviation of the nuclear charge distribution from a spherical shape. It was observed that the electric quadrupole moments for nuclei between closed shells had values much larger than those possible from shell model configurations [Tow49, Gor49]. The discovery of the giant dipole collective mode, where the protons and neutrons oscillate against oue another about the center of mass, greatly motivated further development of collective models.

The first collective descriptions of nuclei modeled the nuclear matter as a classical liquid drop possessing surface deformations from a spherical shape. The simplistic liquid drop model had been developed prior to much of the experimental indications
for collective nuclear behavior as a need to explain the process of nuclear fission. Bohr and Wheeler [Boh39], expanding on ideas of Weizsacker [Wei35], studied shape stability of charged liquid drops deformed from a spherical equilibrium shape. The surface of an arbitrarily deformed liquid drop shape can be constructed by an expansion of spherical harmonics as

$$
\begin{equation*}
R(\theta, \phi)=R_{0}\left[1+\Sigma_{l m} \alpha_{l m} Y_{l m}(\theta, \phi)\right] \tag{1.1}
\end{equation*}
$$

where $R_{0}$ is the mean radius, $\alpha_{l m}$ are deformation parameters, and $Y_{l m}$ are spherical harmonics. This straightforward, but useful description of deformed nuclear shapes makes an explicit counection to the importance of multipole orders. Since the nuclear matter is fairly incompressible, the monopole moment, which represents a compressive "breathing" mode of the uucleus, occurs only at high energies. At modest values of $\alpha_{l m}$, the dipole mode is also somewhat trivial, arisiug from a translation of the nucleus as a whole without a change in the intrinsic nuclear shape. The important leading term in describing collective motion at low energies is, therefore. the quadrupole moment.

Because the quadrupole degree of freedom is the dominant excitation mode at low energy throughout the Periodic Table, most of the various collective pictures have concentrated exclusively on this mode. Although octupole and hexadecapole degrees of freedom have been shown to contribute to the overall stable nuclear shape [Hen68, But96], their contributions are typically much smaller than the dominant quadrupole contribution. Deformation modes of any multipole order are allowed by the expansion given previously in Equation 1.1.

The earliest shell models, too, have been modified and expanded to account for collective behavior. The great success of the shell model in describing the behavior of spherical nuclei near the magic numbers, and the knowledge of collective behavior away from closed shells, led to the development of a deformed shell model. Nilsson's deformed shell model [Nil55] has been widely used with great success in predicting
the ground state spins and parities of many nuclei away from closed shells. It has been particularly useful in describing odd mass nuclei, and its importance can not be overstated.

In the mid-1970's, however, a different approach to describing collective motion was developed. The model, known as the Interacting Boson Model (IBM), is developed from the algebraic properties of the dynamical symmetries of the nucleus [Ari75, Ari76, Ari78, Ari79]. The IBM treats pairs of nucleons coupled to specific angular momentum states as integer spin bosons, and has been useful in describing a wide variety of collective phenomena.

The wide variety of collective behavior is due, in part, to the wide variety of shapes that uuclei can possess. Shape deformations of the uucleus can be either static or dynamic. The simplest example of the excitation of a statically deformed nucleus is rotational behavior. The statically deformed nucleus can be viewed as a classical rigid rotor with energy levels spaced according to a $J(J+1) / 2 I$ rule. where $J$ is the total angular momentum and $I$ is the classical moment of inertia. In contrast, a vibrational excitation is characterized by an energy spectrum which, in the harmonic limit, has a series of equally spaced and degenerate multiplets. The most commonly observed vibrational mode is the quadrupole mode. However, higher order vibrational modes can be excited as well. Octupole vibrations, a $J=3$ multipole order carrying $3 \hbar$ units of angular momentum and hexadecapole ( $J=4$ ) vibrations are found throughout the Periodic Table.

The most commonly observed form of vibrational collectivity is the electric quadrupole excitation [Sch55]. The most obvious signatures of low energy quadrupole vibrations are large transition strengths between energy levels which are nearly evenly spaced, aud the presence of nearly degenerate multiplets which are evenly spaced and correspoud to multi-phonon excitations. Rainwater [Rai50] asserted in 1950 that single valence nucleons outside of a closed shell could polarize
the closed shell core and lead to oscillatory vibrations. New collective degrees of freedom were described by Bohr and Mottelson [Boh52, Boh53, Boh75] which led to the discovery of other vibrational multipoles. Octupole vibrations, evidenced by low energy $J^{\pi}=3^{-}$states, and coupling between quadrupole and octupole degrees of freedom, evidenced by low energy $J^{\pi}=1^{-}$states, were found soon after their important work [Ste54, Nat56, Ken56].

A main focus of experimental curiosity is to examine anomalous behavior in nuclei. Nuclei which exhibit unusual or unexpected behavior allow nuclear structure models to be advanced beyond their simplest forms. Ideally, any useful model will describe many different types of behavior over a wide mass range of nuclei. It is the observed behavior at the very edge of a model's capability which motivates careful modification and adaptation of the model itself.

The present work investigates two different examples of anomalous collective behavior in two different mass regions. The first example is seen in the $74 \leq Z \leq 80$ mass region. This mass region is known informally as the heavy trausitional region due to a shape transition around $A \approx 192$ from prolate to oblate shape, and then to spherical at $Z=8 \%$. Nuclei in the heavy transitional region exhibit a wide range of shapes, from near spherical for the Pb and Hg isotopes with $A \approx 208$, to well-deformed shapes farther from the closed shells. While both the spherical shape and the well-deformed shapes have been well described by the shell model and collective models respectively, the transition region between the two shapes is less understood. This is due, in part, to the excitations being a complex coupling of rotational, vibrational, and single-particle degrees of freedom.

In fact, even some of the pure vibrational modes in this mass region have not been adequately described thus far. In particular, the octupole vibrational mode in the heavy transitional region shows unexpected behavior which warrants the investigation in the present work. In other mass regions, the systematic behavior of
the low energy octupole state is typically smooth as the nucleus undergoes a gradual shape change. The energies of the octupole states vary only slightly from nucleus to nucleus. In contrast to the smooth behavior seen in other mass regions, the energies of the octupole states in the heavy transitional region show abrupt changes.

The second example of anomalous behavior investigated in the present work involves differences between the proton and neutron contributions to a quadrupole vibrational excitation in the $12 \leq A \leq 26$ mass region. Some background of early investigations of proton and neutron differences is necessary here. The Giant Dipole Resonance (GDR) is the earliest example of behavior for which proton and neutron motions are not identical. The GDR is a high energy ( $\geq 15 \mathrm{MeV}$ ) collective excitation. In this mode of excitation, the neutrons and protons oscillate against each other about the center of mass of the nucleus. Since the protons constitute the charge carriers of the nucleus, the GDR is seen as a strong oscillation of charge with respect to the center of mass which gives rise to a large dipole (E1) transition when the mode is excited by electromaguetic probes. Several theories were quickly developed upon discovery of the GDR, the most successful of which were from Goldhaber and Teller [Gol48], Steinwedel [Ste50], and Danos [Dan51]. The various theories formulated the now well known $A^{-1 / 3}$ energy dependence of the GDR, and later extensions of the theory predicted a splitting of the resonance due to the deformation of the uucleus [Oka58, Dau58]. While the GDR is a high euergy collective excitation, and this work is primarily concerned with low energy collective excitations, the GDR is an excellent example of anomalous behavior which prompted further theoretical development.

The GDR also suggests that protons and neutrons can behave differently from one another in a collective motion. That is, the nuclear force can have an "isospin" component. Heisenberg suggested that the proton and the neutron could be treated as two states of the nucleon [Hei32]. Neutrons and protons could then be dis-
tinguished by the projection of the isospin vector $T_{z}$ onto an arbitrarily assigned symmetry axis. With the isospin compouent added to transition operators between nuclear states, differences in proton and ueutron behavior can be quantified on the basis of a nou-vanishing isovector compouent of the transition. Any isovector magnitude is usually treated independently of the known Coulomb charge dependence, which is the dominant isospin symmetry breaking term.

A wonderful test of isospin symmetry is found in the first excited states of eveneven closed shell nuclei in the $12 \leq \mathrm{A} \leq 26$ mass region. The $2_{1}^{+}$states of even-even uuclei in this mass region are often thought of as collective. isoscalar excitations. It has been shown, however, that differences between the proton and neutron contributions can occur in $0_{g s}^{+} \rightarrow 2_{1}^{+}$transitions, especially in single closed shell nuclei [Ber83, Ken92].

A comparison of the neutron and proton oscillation magnitudes can be done by comparing the $B\left(E 2 ; Z_{1}^{+} \rightarrow 0_{g . s .}^{+}\right)$reduced transition matrix element from two different experimental probes. Madsen et al. [Mad75] have stated that a useful comparison of proton and neutron matrix elements can be carried out with low energy ( $10-50 \mathrm{MeV}$ ) proton scattering to measure the neutron contribution, and electromagnetic transitions to measure the proton contribution. While the electromagnetic probe is sensitive to the charge (proton) distribution only, other probes have varying degrees of neutron to proton sensitivity [Ber81]. The neutron and proton contributions to an excited state can be separated, therefore, by using two or more experimental probes.

A particularly useful way of expressing the difference in neutron and proton contributions to the $0_{g s}^{+} \rightarrow 2_{1}^{+}$trausition is by the ratio of the neutron and proton multipole matrix elements $\mathrm{M}_{n} / \mathrm{M}_{p}$. The trausition matrix element for protons (neutrons) between initial and final states is given by

$$
\begin{equation*}
M_{p(n)}\left(T_{z}\right)=<\nu_{f} J_{f} T_{f} T_{Z} A\left\|O_{p(n)}^{\lambda}\right\| \nu_{i} J_{i} T_{i} T_{Z} A> \tag{1.2}
\end{equation*}
$$

where $O_{p(n)}^{\lambda}$ is the multipole transition operator for a transition of order $\lambda$ for protons (neutrons), $J$ is total angular momentum, $T$ is the isospin with projection $T_{Z}=$ $\frac{1}{2}(Z-N)$, and $\nu$ is a generic representation of additional quantum numbers needed to specify the state. For purely isoscalar transitions ( $\Delta T=0$ ), the neutrons and protons oscillate at the same amplitude, and $M_{n} / M_{p}=N / Z$. Deviations result from a non-vanishing isovector contribution to the transition.

A major focus of the present work is to provide further information about the systematics of proton and neutron contributions to low-lying quadrupole excitations in even-even nuclei by a study of the instable nucleus ${ }^{20} \mathrm{O}$. Until receutly, the study of nuclei outside the valley of stability using scattering reactions such as ( $p, p^{\prime}$ ) was not feasible. With recent advances in radivactive beam production and techniques for measuring proton scattering in inverse kinematics, the possibility of studying the more than 3000 particle stable nuclei has been opened to us. Measurements from these least stable nuclei provide stringent tests of any nuclear model, whether collective or single-particle in character.

The present work presents two separate and distinct studies of collective behavior motivated by anomalous systematics. The first study will detail a ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ coincidence experiment performed at Florida State University which populated a second octupole state containing a significant amount of the octupole phonon strength. The $J^{\pi}=3_{2}^{-}$spin assigument given to this state relies on previous proton angular distributions limiting the spin to $J^{\pi}=33^{-}$or $4^{+}$[Cot88b] and previous $\gamma$-ray angular distributions consistent with a $\cdot J^{\pi}=3^{-}$assigument [DiP93].

The second major focus of the present work uffers results from a proton scattering experiment in inverse kinematics from the radioactive beam ${ }^{20} \mathrm{O}$ performed at the Natioual Superconducting (yyclotron Laboratory. The proton angular distribution is used to extract an effective quadrupole deformation parameter $\beta_{2}$, which, in conjunction with the known $B\left(E 2 ; 0\right.$ g.s. $\left.\rightarrow 2_{1}^{+}\right)$electromagnetic measurement
[Ram87, Ti195], yields a value of $M_{n} / M_{p}=2.8 \pm 0.4$. The result is much larger than $N / Z$, which is expected for an isoscalar transition, and clearly indicates an enhanced neutron contribution to the transition.

Chapter 2 of the present work is a theoretical overview of octupole excitations in the heavy transitional mass region. Experimental details of the ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ coincidence experiment are given in Chapter 3 , followed by results from the experiment and a discussion of the results in Chapters 4 and 5. Cbapter 6 is a theoretical overview of $2_{1}^{+}$states in the $12 \leq A \leq 26$ mass region. Chapter 7 presents the experimental details of the inverse kinematics $p\left({ }^{20} \mathrm{O}, p\right)^{20} \mathrm{O}^{*}$ proton scattering experiment. The analysis and results from the experiment are given in Chapter 8 , followed by a discussion of the results in the context of systematics of the mass region in Chapter 9. Chapter 10 will then provide a brief summary of the major points from both experiments.

## CHAPTER 2

## THEORETICAL OVERVIEW OF OCTUPOLE STATES IN ${ }^{196} \mathbf{P t}$

Nuclear structure is typically studied within the context of two major frameworks: the shell model and the collective model. As was mentioned in Chapter 1, the bulk of nuclear behavior exhibits properties of both of these two idealized pictures.

### 2.1 Structure of Excited States in $74 \leq Z \leq 80$ Mass Region

In the shell model, uucleons in an even-even nucleus in the ground state successively fill nuclear orbitals of lowest available energy. In a $j-j$ coupling scheme, pairs of like nucleons couple to zero orbital angular momentum states. Each angular momentum state $j$ is split by coupling to the magnetic moment of the nucleon into a series of magnetic substates $m=-j,-j+1, \ldots, j-1, j$. Each $m$ substate contains oue nucleon, and substates of opposite $m$ sigu couple to zero, giving an overall contribution of zero to the angular momentum of the entire nucleus. Excited states are built by successively promoting nucleons from the ground state configuration to other shells. The angular momentum of the excited state results from the coupling of the excited nucleons with each other and the "holes" left behind in the ground state configuration.

The collective treatment, by contrast, makes no explicit reference to individual nucleon orbits. Instead, eigenstates are built from linear combinations of spherical harmonics which describe the radius of the nuclear surface, as in Equation 1.1, by multipole order. Collective excitations are generally identified by transition
strengths which are considerably larger than those which occur in single-particle transitions. Since the large transition strengths are many times those obtainable with single-particle excitations, we can consider the collective excitation to be a coherent sum of single-particle excitations [Bro59]. The transition probability for an electric transition of multipole order $\lambda$ between two states is related to the reduced transition probability, $B(E \lambda ; i \rightarrow f)$, by

$$
\begin{equation*}
T(E \lambda)=\alpha c \frac{8 \pi(\lambda+1)}{\lambda[(2 \lambda+1)!!]^{2}}\left(\frac{E_{\gamma}}{\hbar c}\right)^{2 \lambda+1} B(E \lambda ; i \rightarrow f) \tag{2.1}
\end{equation*}
$$

with the energy difference between the initial and final states given by $E \gamma . B(E \lambda)$ relates to the full transition matrix elements by

$$
\begin{align*}
B(E \lambda ; i \rightarrow f) & \left.=\sum_{\mu M_{f}}\left|\left\langle\nu_{f} J_{f} M_{f}\right| O_{\lambda \mu}^{E}\right| \nu_{i} J_{i} M_{i}\right\rangle\left.\right|^{2}  \tag{2.2}\\
& \left.=\frac{1}{2 . J+1}\left|\left\langle\nu_{f} J_{f} \| O_{\lambda}^{E}\right|\right| \nu_{i} J_{i}\right\rangle\left.\right|^{2}
\end{align*}
$$

where the $\nu$ indicates additional quantum numbers needed to uniquely determine the state, and $O_{\lambda \mu}^{E}$ is the electric multipole operator given by

$$
\begin{equation*}
O_{\lambda_{\mu}}^{E}=\sum_{i} r_{i}^{\lambda} Y_{\lambda \mu}\left(\hat{r}_{i}\right) . \tag{2.3}
\end{equation*}
$$

The Weisskopf unit (W.u.) single-particle estimate [Blai9], is commonly used to quantify the strength of a trausition independently of the size of the nucleus. This is done by considering the trausition probability of multipole order $\lambda$ of a single nucleon moving from one single-particle orbit to another. In this extreme independentparticle picture, the transition of a single particle is assumed to occur without affecting any other nucleons. The single-particle estimate can be obtained from the transition operator by evaluating the average value of $r^{\lambda}$ by

$$
\begin{equation*}
\left\langle r^{\lambda}\right\rangle=\int_{0}^{\infty} R_{f}^{-} r^{\lambda} R_{i} r^{2} d r \tag{2.4}
\end{equation*}
$$

where $R_{f}^{*}$ and $R_{i}$ are the normalized radial parts respectively of the final and initial single-particle wavefunctions. Approximating the nucleus as a uniformly dense sphere with radius $R=r_{o}^{\lambda} A^{\frac{\lambda}{3}}$ fm reduces the the radial integral to

$$
\begin{equation*}
\int_{0}^{\infty} R_{f}^{*} r^{\lambda} R_{i} r^{2} d r \approx \frac{3}{\lambda+3} r_{o}^{\lambda} A^{\frac{\lambda}{3}} . \tag{2.5}
\end{equation*}
$$

The Weisskopf single-particle estimate for electric transitions is then given by

$$
\begin{equation*}
B_{W}(E \lambda)=\frac{1}{4 \pi}\left(\frac{3}{\lambda+3}\right)^{2} r_{o}^{2 \lambda} A^{2 \lambda / 3} e^{2} f m^{2 \lambda} \tag{2.6}
\end{equation*}
$$

### 2.2 Interacting Boson Model

The behavior of the octupole states in the even-even Pt isotopes has been discussed in the framework of the Interacting Boson Model (IBM) [Zam93a], the Quasiparticle-Phonou Model (QPM) [Pon92], and Hartree-Fock formalisms [Egi96]. Only the IBM and QPM approaches, however, have been successful in explaining fragmentation of the octupole strength.

Until the mid-1970s, descriptions of collective behavior were generally given in the Bohr and Mottelson geometric picture. An alternate method of describing collective behavior, the IBM, was introduced by Arima aud Iachello in 1974 [Iac 74 , Ari75], and is algebraic in its formalism. The IBM is an attempt at truncating the shell model basis space which quickly becomes too large to be useful in describing the behavior of heavy nuclei. The IBM treats valence uucleons in pairs as bosons which couple to particular integer spins. Low-lying collective behavior is then described in terms of the energies and interactions of these nucleon pairs. The truncation of the shell model space results from allowing only a limited number of angular momentum values $J$ to which a pair of nucleous can couple. For example, if momentum couplings of only $J=0$ ( $s$ bosons) and $J=2$ ( $d$ bosons) are allowed, the number of magnetic substates is limited to six ( 1 from $J=0,5$ from $J=2$ ), and basis states
span a six dimensional space. The allowed angular momentum couplings constitute a system which can be described in terms of its $U(6)$ algebraic group properties.

The $U(6)$ group can be decomposed into subgroups by way of three decomposition chains which are physically meaningful. That is, the three chains decompose to an $O(3)$ symmetry which is required for invariance under rotation. Each of the three chains has a geometric aualog associated with it. Consequently, the three group chains define the limiting cases of collective behavior associated with the quadrupole degree of freedom in the IBM picture. The $U(5)$ symmetry is the IBM equivalent of a spherical vibrator. $\mathrm{SU}(3)$ symmetry corresponds to an axially symmetric rigid rotor deformation commonly assuciated with rotational collective motion. The $\mathrm{O}(6)$ symmetry, which is the primary concern of this section, corresponds to a $\gamma$-unstable, axially-asymmetric rotor [Ciz78, Cas78, Bol81].

Some rather bold and drastic assumptious are made in the IBM. Valence uucleons make the primary contribution to the excitation, and excitations from uucleons in closed shells are taken into account by means of effective charges in the transition operators. Additionally, no distinction is made between protons and neutrons, in the earliest version of the model (IBM-1), nor is there a distinction between particles and holes. Valence nucleons are counted from the closest closed shell. Some of these restrictions to the model are removed in advanced forms of the IBM (IBM-2 and IBM-3). Casten and Warner provide an excellent review of the IBM formalism [War82]. The present work will be concerned ouly with the usage of the IBM in understanding octupole behavior. Some background is necessary here, though, to make this possible.

The IBM features a Hamiltoniau containing operators for the creation and annihilation of the bosons needed to construct the excitations of interest. A generic operator is written in the spherical tensor form used by Barfield et al. [Bar88] $\tilde{b}_{\mu}^{(\lambda)}=(-1)^{(\lambda+\mu)} b_{-\mu}^{(\lambda)}$. In the simple case of a harmonic oscillator, for example, the

Hamiltonian is given by

$$
\begin{equation*}
H=\epsilon_{d}\left(d^{\dagger} \tilde{d}\right)=\epsilon_{d} n_{d} \tag{2.7}
\end{equation*}
$$

where $\epsilon_{d}$ is the energy of a $d(J=2)$ boson, and $d^{\dagger}$ and $d$ are the creation and annhilation operators respectively of a $d$ boson. The energy of the excitation is then just the intrinsic energy of the $d$ boson multiplied by the number $n_{d}$ of $d$ bosons needed to create the excitation.

In order to describe octupole excitations, however, the addition of an $f(\mathrm{~J}=3)$ negative parity boson is required. A suitable Hamiltonian for describing a number $N_{B}$ of bosons in an $(s d)^{N_{B}-x} f^{x} \mathrm{O}(6)$ basis, where $x=0$ or 1 , is given by

$$
\begin{equation*}
H=H_{s d}+H_{f}+V_{s d f} . \tag{2.8}
\end{equation*}
$$

This choice of basis allows for, at most, a single $f$ boson. The first term of the Hamiltonian, $H_{s d}$, describes the even-even positive parity ( $s d$ ) core upon which the negative parity excitations are built, described by $H_{f}$. The final term, $V_{s d f}$, is a two-body interaction between the positive and negative parity bosons. The positive parity core term consists of pure $d$ boson operators, $L_{d}$, and a mixing operator between $s$ and $d$ bosons, $Q_{d}$, in the form

$$
\begin{equation*}
H_{s d}=\epsilon n_{d}+a_{1} L_{d} \bullet L_{d}-a_{2} Q_{d} \bullet Q_{d} \tag{2.9}
\end{equation*}
$$

where $n_{d}=d^{\dagger} d$ is the $d$ boson number operator and

$$
\begin{align*}
L_{d} & =\sqrt{10}\left(d^{\dagger} \tilde{d}\right)^{(1)}  \tag{2.10}\\
Q_{d} & =\left(s^{\dagger} \tilde{d}+d^{\dagger} s\right)+\chi_{2}\left(d^{\dagger} \tilde{d}\right)^{(2)} \tag{2.11}
\end{align*}
$$

The transition from $O(6) \rightarrow \mathrm{SU}(3)$ symmetries is obtained by varying the free parameter $\chi_{2}$ from $\chi_{2}=0$ (the ideal $\mathrm{O}(6)$ limit) to $\chi_{2}=-1.32$ (the $\mathrm{SU}(3)$ limit). The $f$ boson excitations are handled by a simple

$$
\begin{equation*}
H_{f}=\epsilon_{f} n_{f} \tag{2.12}
\end{equation*}
$$

where $n_{f}=f^{\dagger} \tilde{f}$, the $f$ boson number operator, equals zero for positive parity excitations and one for negative parity excitations in this choice of basis which is limited to a single $f$ boson. The interaction between the positive parity core and the $f$ boson system is cousidered as a multipole expansion of the form

$$
\begin{equation*}
V_{s d f}=A_{1} L_{d} \bullet L_{f}+A_{2} Q_{d d} \bullet Q_{f}+A_{3}: E_{d f}^{\dagger} \bullet E_{d f}: \tag{2.13}
\end{equation*}
$$

with

$$
\begin{align*}
L_{f} & =2 \sqrt{7}\left(f^{\dagger} \tilde{f}\right)^{(1)}  \tag{2.14}\\
Q_{f} & =-2 \sqrt{7}\left(f^{\dagger} \tilde{f}\right)^{(2)}  \tag{2.15}\\
: E_{d f}^{\dagger} \bullet E_{d f}: & =5:\left(d^{\dagger} \tilde{f}\right)^{(3)} \bullet\left(f^{\dagger} \tilde{d}\right)^{(3)} \tag{2.16}
\end{align*}
$$

The : $E_{d f}^{\dagger} \bullet E_{d f}$ term is an "exchange" term composed of a linear combination of five multipole terms, and is added for phenomenological reasons [Bar88].

### 2.3 Octupole Fragmentation in the Heavy Transitional Region

It has long been known that the low energy octupole state is fragmented in well deformed nuclei, such as those in the rare earth and actinide regions [ Nee 70 , $\operatorname{Cot} 96$ ]. In the case of nuclei with large static quadrupole deformations, the octupole mode aligns along different $K$ quantum number components with respect to the symmetry axis having $K=0,1,2$, and 3 . The Pt isotopes of the heavy transitional region, however, do not have large, stable deformations, having $\beta_{2} \leq 0.15$. The high degree of fragmentation observed in the Pt isotopes can not be explained simply by different $K$ alignments.

As early as 1986 Engel [Eng86] predicted fragmentation of the low energy octupole state in the Pt isotopes. Zamfir et al. [Zam93a] further demonstrated that octupole fragmentation would occur in nuclei which conform to the $\mathrm{O}(6)$ dynamical
symmetry, of which ${ }^{196} \mathrm{Pt}$ is the best known example [Ciz78]. Zamfir et al. summarized two different situations where octupole fragmentation is likely to occur in the IBM formalism. The first is seen in nuclei having a large static quadrupole deformation. The nuclei in this category correspond to the usual picture where large fragmentation is expected, and are interpreted in the IBM as a transition from $U(5)$ to $S U(3)$ symmetry. The other situation where large octupole fragmentation is expected is for nuclei conforming to an $\mathrm{O}(6)$ symmetry, such as the Pt isotopes. As mentioned previously, $O(6)$ nuclei have $\gamma$-soft, axially asymmetric shapes and do not have large static deformatious [Ciz78, Cas78, Bol81]. It has even been proposed that the existence of octupole fragmentation can provide a signature for $\mathrm{O}(6)$ nuclei [Zam93a].

The calculations of Zamfir ct al. [Zam93a, Zam9:3b] show that a high degree of fragmentation of the low energy octupole state is expected for the even-even Pt isotopes. In their analysis, a large concentration of octupole strength is found nearly 1 MeV above the $3_{1}^{-}$state. Distributions of octupole strength from their calculations are given in Figure 2.1 and show the additional strength contained in higher-lying octupole states. The example used in the figure is from ${ }^{194} \mathrm{Pt}$, but nearly identical results are obtained for ${ }^{196,198} \mathrm{Pt}$ as well. Zamfir et al. also demonstrate that the large degree of fragmentation observed in the $P t$ isotopes cannot be reproduced by modeling the region as a $U(5) \rightarrow S U(3)$ transition region. Only by modeling the region with an $O(6)$ symmetry are they able to reproduce the experimentally observed magnitude of the octupole fragmentation in the IBM-1 formalism.

Zamfir et al. made additional simplifications to the IBM Hamiltonian of Equation 2.8 in calculating octupole fragmentation for a wide range of nuclei [Zam93a]. In particular, $\epsilon_{f}$, the overall excitation energy of the negative parity states, is set to unity. Both $a_{1}$ of Equation 2.9 and $A_{1}$ of Equation 2.13 are set to zero because the $L_{d} \bullet L_{d}$ and $L_{d} \bullet L_{f}$ terms have little effect on the non-yrast $3^{-}$states wavefunctions,
and therefore little effect on the magnitude of the fragmentation. Additionally, $a_{2}=0.05 \mathrm{MeV}$ was used for all the calculations, and $\chi_{2}$ was varied between values of $0.22-0.055 N_{B}$. Zamfir et al. used the same IBM parameters that were used in an earlier study of the positive parity core [Zam9:3b].

No new IBM calculations were performed for the present analysis, which focuses on identification of low-lying octupole states in the even-even Pt isotopes. The preceding section is intended ouly to provide a reasonable framework for interpreting the Pt isotopes as examples of the $\mathrm{O}(6)$ symmetry in which a high degree of fragmentation is expected for the low energy octupole state.


Figure 2.1: Octupole state strength distribution. (a) ${ }^{194} \mathrm{Pt}$ modeled as an $\mathrm{O}(6)$ nucleus, (b) ${ }^{146} \mathrm{Nd}$ modeled as a U(5) nucleus, and (c) ${ }^{168} \mathrm{Er}$ modeled as an $\mathrm{SU}(3)$ nucleus. Of the three idealized symmetries, only the $O$ (6) symmetry can reproduce the significant octupole strength in higher-lying $3^{-}$states in the IBM-1 formalism.

## CHAPTER 3

## ${ }^{196} \operatorname{Pt}\left(p, p^{\prime} \gamma\right)$ COINCIDENCE EXPERIMENT

We performed a proton - $\gamma$ coincidence experiment at the Florida State University Tandem-LINAC facility to study low-lying octupole states in ${ }^{196} \mathrm{Pt}$. A $2.1 \mathrm{mg} / \mathrm{cm}^{2}$ thick self-supporting target of $99 \%$ isotopically euriched ${ }^{196} \mathrm{Pt}$ was bombarded by a beam of 12.7 MeV protous. Both elastically and inelastically scattered protons were detected, as were the $\gamma$-ray de-excitations from excited target nuclei.

### 3.1 Beam Production

A beam of 12.7 MeV protous was produced from a commercially bought hydrogen cathode placed in a SNICS (source of negative ions by cesium sputtering) source. The SNICS source produces negative ions as an atomic beam from oven-heated cesium. The cesium beam is ionized as it contacts a hot, coaxial tantalum surface. and is accelerated toward a negatively biased cathode material from which neutral hydrogen atoms are sputtered. The hydrogen atoms pick up electrons as they pass through a layer of cesium condensed on the front of the cathode. The negatively charged hydrogen ions are then injected into the tandem Van de Graaff accelerator. Source intensities of up to $3 \mu A$ of negatively charged protons are obtainable.

The proton beam was accelerated by the Florida State Uuiversity FN tandem Van de Graaff to an energy of 12.7 MeV . The negatively charged beam was stripped of electrons at the tandem terminal to produce a positively charged proton beam. The beam was focused by a series of quadrupole magnets and steered by dipole magnets into the experimeutal vault. We choose a beam energy of 12.7 MeV on
the basis of a number of ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ measurements performed by the author. A range of beam energies between 10 MeV and 17 MeV were tried before deciding on 12.7 MeV . The lower limit of 10 MeV represents the minimum energy required to penetrate the Coulomb barrier, while 17 MeV was the maximum energy that did not produce so many neutrons as to overwhelm the detectors. In each trial, the total yields of elastic and inelastic proton events was monitored, as were background ueutron and $\gamma$-ray radiation yields. The 12.7 MeV beam energy appears to optimize the inelastic scattering cross section relative to the elastic, while causing minimal $\gamma$-ray background from ( $p, n$ ) and ( $p, 2 n$ ) reactions. The energy is comparable to that used in a similar ${ }^{198} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ study [YatS8 8 ].

### 3.2 Particle - $\gamma$ Detection

A new beam line and scattering chamber were constructed in the old target room for the particle - $\gamma$ experiment. The new beam line was necessary to position the scattering chamber as close to the final quadrupole focusing magnet as possible to obtain a small, well-focused beam spot on the target. It is crucial in proton - $\gamma$ experiments of this kind to minimize any background neutron radiation. The neutrons not only damage the $\gamma$-ray detectors. but also give rise to a high $\gamma$-ray background. Therefore, it is imperative that the beam strikes ouly the target material and does not impinge upon the target frame. Similarly, collimators could not be used to reduce the beam size because unwanted neutrou and $\gamma$-ray background from protons striking the collimators would have resulted. 12.7 MeV protons are above the Coloumb barrier of most common collimator materials, and thus react readily with the collimators themselves. Since the beam spreads as a function of distance from the quadrupole magnet, it is advantageous to place the target at the minimum focal leugth of the magnet, thereby eliminating the need for collimators.

The beam dump consists of a 4.6 m long section of 15 cm beam pipe. The large diameter pipe and the large length of the beam dump is required to reduce background radiation. The end of the beam dump contains a natural carbon disk placed in the beam path. To further reduce the neutron background, the last 1.2 m of the beam dump is surrounded with 15 cm of boronated paraffin shielding. The boron has a large cross-section for absorbing neutrons, and the paraffin acts a moderator to slow the neutrons. The nentron dose rate along the beam line was reduced from $2000 \mathrm{mrem} / \mathrm{hr}$ during the first test run, to $2 \mathrm{mrem} / \mathrm{hr}$ during the production runs by this shielding and by carefully tuning the incident beam.

The scattering chamber built for the present experiment is a 20 cm diameter cylindrical chamber large enough to house a charged-particle detector and its cooling apparatus. The chamber radius is kept at a minimal size to allow $\gamma$-ray detectors to be placed as close as possible to the target, thereby increasing the target event to background ratio. In an effort to minimize $\gamma$-ray attenuation, the chamber walls were constructed at a minimal thickness needed to maintain structural integrity under vacuum. The chamber allows up to four targets to be inserted into the beam path without breaking vacuum. Figure 3.1 is a schematic drawing of the detector positions.

A single charged-particle detector was used to detect scattered protous. The detector was a $100 \mathrm{~mm}^{2}$ area silicon surface barrier detector, having a thickness of 1.5 mm and placed 6 mm from the target position. The detector was thick enough to fully stop protous up to 20 MeV . The charged-particle detector was placed at $90^{\circ}$ to the incident beam direction, an angle at which the elastically scattered proton cross section is much smaller than that at more forward angles. The cross section for inelastically scattered protons at $90^{\circ}$ is comparable to that at more forward angles.

The particle detector has an intrinsic resolution of 50 keV for detecting alpha particles from ${ }^{228} \mathrm{Th}$ and ${ }^{241} \mathrm{Am}$ radioactive sources. The in-beam resolution, how-


Figure 3.1: Schematic of the ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ experimental apparatus.
ever, was considerably worse ( 150 keV ) due to "pile up", a successive particle event occurring in the detector before the charge from the preceding event has been collected. A major contributor to the pile up was beta decay from the activated target in which ${ }^{196} \mathrm{Au}$ had been produced via the ${ }^{196} \mathrm{Pt}(p, n)$ reaction. After a few hours of beam on target, the beta decays became frequent enough to interfere with the detector resolution. A number of ways to reduce the resolution degradation were investigated, including placing magnets around the detector entrance path to deflect unwanted electrons, and biasing the target with positive voltage to capture the produced electrous. These methods had only a small effect on the pile up.

A compromise was made, finally, between maximizing the beam current on target and minimizing the resolution of the charged particle detector. During the production runs, the detector was cooled by thermal contact with a brass block in which ethanol chilled to $-20^{\circ} \mathrm{C}$ was circulated. Cooling the detector minimizes the effects of neutron damage in the detector.

Detecting $\gamma$-ray de-excitations was considerably more straight forward than particle detection. Four Compton-suppressed n-type hyper-pure Ge (HPGe) detectors having $20 \%$ relative efficiency and 2.5 keV resolution at 1.33 MeV from the FSU-Pitt Gamma Array [Tab93] were used to identify $\gamma$-ray de-excitations from excited target nuclei. All four detectors were placed at $90^{\circ}$ to the beam direction in a semi-circle surrounding the scattering chamber (see Figure 3.1). The detectors were placed as close as possible ( 6 cm ) to the target to maximize the beam event to background ratio. Even so, the $\gamma$-ray background was significant.

The $\gamma$-ray background is mainly produced in two ways. The first is from reactions away from the target position. As mentioned previously, 12.7 MeV protons are above the Coulomb barrier on most materials commonly used in beam line construction. Neutrons and $\gamma$-rays were produced, therefore, around the quadrupole focusing maguet before the detector, and from secondary scattered protons after the target. Although Compton suppression of the Ge detectors reduces this background to a great extent, a large amount of shielding and very careful beam tuning was required to further reduce the background. The beam tuning was performed while monitoring radiation flux along the beam line. The second major source of $\gamma$-ray background came from ( $p, n$ ) and ( $p, 2 \mathrm{n}$ ) reactions of protons with the ${ }^{196} \mathrm{Pt}$ target. The dominance of ( $p, \mathrm{n}$ ) and ( $\mathrm{p}, 2 \mathrm{n}$ ) reactions was also seen in ${ }^{198} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ experiments [Yat88]. Unlike the first source, these background events take place at the target position and can not be reduced by shielding. Fortunately, $\gamma$-rays from $(p, n)$ and ( $p, 2 n$ ) reactions are not associated with a scattered proton, and a care-
ful particle- $\gamma$ coincidence requirement can filter out a majority of these unwanted events. Even with a coincidence requirement, however, many "chance" coincidence events were seen between an elastically scattered proton and a $\gamma$-ray de-excitation from ( $p, n$ ) and ( $p, 2 n$ ) reactious.

The particle- $\gamma$ coincidence requirement is crucial for two reasons. The first is to reduce the dominating background from ( $p, n$ ) and ( $p, 2 n$ ) reactions just mentioned. The second, and perhaps more important reason, is to associate each $\gamma$-ray decay with a particular excitation of the target nucleus. The particle- $\gamma$ coincidence condition required a particle event followed by a $\gamma$-ray event from any one of the four Ge detectors within a 200 us time window. To further separate true coincidences from the chance coincidences falling within the 200 ns time window. the time difference between the detected particle and the $\gamma$-ray in coincidence with it was recorded via a time-to-amplitude converter (TAC). The TAC' was started by a particle event and stopped by any $\gamma$-ray event. Prompt coincidences resulting from an inelastically scattered proton and the $\gamma$-ray de-excitation of the populated state all occur at one specific time difference, while chance coincidences from other $\gamma$-ray de-excitations occur over a continuum of time differences within the 200 us time window. The TAC spectrum (Figure 3.2 (a)) shows a peak at the prompt coincidence time difference sitting on top of the chance coincidence events. The prompt TAC peak was used as a gating requirement in the off-line analysis.

The collected particle- $\gamma$ events were sorted in two separate matrices of particle energy versus $\gamma$-ray energy. One of the matrices was sorted with a gate on the prompt time peak seen in the TAC spectrum, and the other was sorted with a gate on an equal number of chanuels of the chance coincidence event time background. The chance time gated matrix was subtracted from the prompt time gated matrix to produce a single time background subtracted particle- $\gamma$ matrix. Figures 3.2 (b) and (c) show the dramatic effect of using the TAC to filter chance coincidences. The


Figure 3.2: The TAC timing spectrum and gated particle projections. (a) The time spectrum is the total yield as a function of the time difference between a particle and $\gamma$-ray event. A peak from the prompt coincidences sits above a continuum of "chance coincidence" background. (b) Projected proton spectrum before any software gating. The first excited state is seen as a shoulder of the elastic proton peak. (c) Projected proton spectrum with the chance coincidences subtracted. The elastic peak is nearly completely suppressed, and the spectrum is dominated by the first excited state.
proton spectrum in Figure 3.2 (b) is shown without any TAC gating condition. The first excited state in ${ }^{196} \mathrm{Pt}$ appears as a "shoulder" on the dominant elastic peak. The same proton spectrum is shown again with the chance coincidences subtracted from the prompt in Figure 3.2 (c). The elastic peak, which dominated the spectrum when no gating was imposed is nearly completely suppressed, while the first excited state now dominates the spectrum.

### 3.3 Particle- $\gamma$ Electronics

The electronics requirements were fairly straightforward for the experiment. Figure 3.3 is a diagram of both the signal and logic electronics. The particle signals were split after pre-amplification into energy and timing signals. The energy signal was fed directly into a spectroscopy amplifier for signal amplification and shaping. The timing signal goes first to a fast timing filter amplifier (TFA) for sigual shaping. The output is fed to a constant fraction discriminator (CFD) where the leading edge of the signal is picked off and used to start the TAC.

The Ge detector electrouics were set up in a similar fashion. The energy signal was pre-amplified and sent directly to a spectroscopy amplifier. The timing signal from the preamplifier feeds a TFA, whose output is discriminated by a ('FD. The CFD output goes to a "blue box" coincidence unit to be used in Compton suppression and to set a logical OR condition for the four Ge detectors. In its capacity as a logical OR unit, any timing sigual from one of the four Ge detectors generates a new timing signal from the coincidence unit.

Compton suppression is haudled by the same "blue box" coincidence unit built at Florida State University specifically for that purpose. The unit allows for easy adjusting of the delay and width of both the Compton shield and the Ge detector timing signals. The Compton BGO shield signal is fed to a TFA, whose output is picked off by a CFD. The CFD output feeds the "blue box" unit, where internal
timing signals are generated. The shield timing signal is adjusted in time to fit completely around the (ie detector signal, and is used to veto any Ge signal for which a shield event is seen. The veto is made internally so that a timing signal is generated only for nou-Compton scattered events. A coincidence requirement set to a multiplicity of one is used to logically OR the four Ge signals. Therefore, any non-Compton scattered event from any of the four Ge detectors results in a positive timing signal. The timing signal is delayed by 280 us and used as a stop condition for the TAC:

Coincidence gating by the TAC is done by using the TAC output, which signifies a particle event followed by a $\gamma$-ray event within 200 ns , to generate a positive particle- $\gamma$ coincidence signal. The signal is lengthened in time by a gate and delay generator (GDG) to fit completely around the spectroscopy amplifier energy signals from all detectors, and is used as a master gate at the analog-to-digital converters (ADCs) for data acquisition.


Figure 3.3: Diagram of the electronics. Both the signal and logic electronics are included.

## CHAPTER 4

## ANALYSIS AND RESULTS OF ${ }^{196}$ Pt DATA

### 4.1 Previous Experimental Work on ${ }^{196} \mathbf{P t}$

Previous experimental [Cot88a, Cot88b, Yat88, Zam93a] work shows that the lowest euergy octupole states ( $3_{1}^{-}$states) in the even-even Pt isotopes ( ${ }^{194.196 .198} \mathrm{Pt}$ ) and some of the Os isotopes have substantially lower excitation energies than do the corresponding states in the rest of the heavy transitional mass region. The lowest energy octupole states in the stable eveu-eveu Hg and Pb nuclei are around 2.6 MeV . This value remains fairly constant throughout the mass region as evidenced by Figure 4.1. The figure depicts the euergy of the lowest-lying octupole state, the $3_{1}^{-}$state, for nuclei in the heavy transitional mass region. The energy for the lowest octupole state in the Pt isotopes, represented as triangles in Figure 4.1, resides nearly 1 MeV lower than in the other nuclei of the region. In addition, the drop in the energy of this $3_{1}^{-}$state is seen to be rather abrupt. This abrupt drop is not seen elsewhere in the heavy transitional region, nor is it seen in other mass regions above $\mathrm{A} \geq 16$ [Spe88].

Cottle et al. suggested that the anomalous lowering of the $3_{1}^{-}$state energy in the even-even Pt isotopes was due to a fragmentation of the octupole phonon strength [Cot88a]. They proposed that the octupole phonou strength was not concentrated in the $3_{1}^{-}$state, but rather was shared between a number of $3^{-}$states. The identification of additional $3^{-}$states is a major focus of the present work and will be discussed later. If a number of additional low energy octupole states could be identified, then a better measure of the octupole state energy may be the energy of the centroid,


Figure 4.1: Energy systematics of the $3_{1}^{-}$states in the heavy transitional mass region ( $74 \leq \mathrm{Z} \leq 82$ ). The energies of the lowest octupole state in Hg and Pb are nearly constant around 2.6 MeV . In the Pt and Os isotopes, however, (shown respectively as triangles and inverted triaugles) the euergy of the lowest octupole state is nearly 1 MeV lower.
defined as

$$
\begin{equation*}
C=\frac{\sum_{i} E_{i} B\left(E 3 ; 0_{g . s}^{+} \rightarrow 3_{i}^{-}\right)}{\sum_{i} B\left(E 3 ; 0_{g . s}^{+} \rightarrow 3_{i}^{-}\right)} \tag{4.1}
\end{equation*}
$$

The centroid measure is a weighted average of the octupole state energy. The energy of each octupole state is scaled by the amount of the octupole phonon strength $\left(B\left(E 3 ; 0_{g s}^{+} \rightarrow 3^{-}\right)\right)$it carries. Looking again at the mass region systematics of the octupole state energy in Figure 4.2 , we see that the value of the octupole state energy for the Pt isotopes, indicated by the open triangles, is very similar to those in Hg and Pb when using the centroid measure. If, in fact, the octupole state strength is fragmented over a number of states, then fragmentation provides an explanation for the observed lowering of the $3_{1}^{-}$state energy. As a simple measure of the magnitude of the octupole fragmentation, Cottle et al. offered the quantity

$$
\begin{equation*}
\Delta E 3=C-E\left(3_{1}^{-}\right) \tag{4.2}
\end{equation*}
$$

where $C$ is the centroid of the octupole state energy as defined previously, and $E\left(3_{1}^{-}\right)$ is the energy of the lowest octupole state. It is seen, then, that if the octupole strength is concentrated in a single state, the degree of fragmentation is zero. As the strength spreads out over a large energy range, the fragmentation increases, as seen in the Pt isotopes. The proposal of octupole fragmentation provides a nice explanation for the reduced euergy of the $3_{1}^{-}$state, and is understood in the eveneven Pt isotopes as an expression of $\mathrm{O}(6)$ symmetry in the IBM formalism [Zam93b].

Octupole fragmentation has been observed before. In fact, large degrees of fragmentation of vibrational phonons have been observed and are well understood for many well-deformed nuclei [Nee70, Bar88, Cot90]. However, the even-even Pt isotopes of the heavy transitional region are not well-deformed. Also, there is no experimental example which indicates that this large a fragmentation of the octupole phonon strength is expected. Indeed, a fragmentation of this magnitude can not be


Figure 4.2: Energies of the low energy octupole state in the heavy transitional region. The open shapes indicate the centroid energy of the low energy octupole state, while the closed shapes are the $3_{1}^{-}$state energies.
seen in any other mass region [Cot90]. The ratio

$$
\begin{equation*}
R_{4 / 2}=\frac{E\left(4_{1}^{+}\right)}{E\left(2_{1}^{+}\right)} \tag{4.3}
\end{equation*}
$$

provides a measure of deformation which allows a study of how the octupole deformation depends on quadrupole deformation. The magnitude of octupole fragmentation seen in the Pt isotopes can be compared in a meaningful way to other nuclei. $\mathrm{R}_{4 / 2}$ has long been used as a structural indicator of the degree of quadrupole deformation for medium and heavy mass unclei. Values of 2.0 indicate a spherical harmonic vibrator, while values of 3.3 represent au upper limit to the most deformed nuclei known. The Pt isotopes in question have $R_{4 / 2} \approx 2.3$, indicating that they do not possess a large static deformation. Looking at a plot of $\Delta E 3$, the degree of octupole fragmentation, as a function of $R_{4 / 2}$ in Figure 4.3, shows the extremely high degree of fragmentation seen in the eveu-even Pt isotopes. They show a larger degree of fragmentation than even the most statically deformed nuclei. The present work seeks to confirm the occurrence of octupole fragmentation in the even-even Pt isotopes by confirming $J^{\pi}=3^{-}$assignments for the presumed octupole states.

The first step in understanding the octupole fragmentation is to identify higherlying $3^{-}$states in the even-even Pt isotopes. A considerable amount of information and study has been done previously which lends insight to the present work. In 1981 Deason et al. [Dea81] performed proton scattering experiments at a 35 MeV beam energy on ${ }^{194,196,198} \mathrm{Pt}$. A number of states in each Pt isotope were populated in the reaction which were identified as good candidates of higher-lying octupole states. Deason et al. made tentative spin assignments of the same states based on their measurements of angular distributions from scattered protons. A later proton- $\gamma$ coincidence experiment performed by Yates et al. [Yats8] on ${ }^{198} \mathrm{Pt}$ identified a 2603 keV state which was tentatively assigned $J^{\pi}=3^{-}$. Yates et al. further noted that the 2603 keV state corresponded to one of the states seen in the ( $p, p^{\prime}$ ) experiment of Deason et al.. Prompted by this new development, a further analysis of the


Figure 4.3: The magnitude of fragmentation is plotted against $R_{4 / 2}$, a structural indicator of the degree of static quadrupole deformation. The degree of fragmentation seen for the Pt isotopes is greater than that of even the most statically deformed examples of nuclei.
scattered proton angular distributions was performed by Cottle et al. [Cot88a] to more definitely determine the spins and parities of some of the higher-lying states. Cottle et al. modeled the proton angular distributions with a coupled channels analysis and compared the angular distributions to lower-lying states with known spin. A total of 11 states in the three Pt isotopes ( ${ }^{194,196,198} \mathrm{Pt}$ ) were identified as being higher-lying octupole states containing a significant amount of the total octupole phonou strength. However, uncertainty existed in the spin assignments of these states. It is clear by looking at the example in Figure 4.4 that the proton angular distribution of a $4^{+}$state is nearly identical to that of a $3^{-}$state in the data set of Deason ct al.. Hence, a $4^{+}$spin assignment could not be excluded for the suggested octupole states in question.

For one of the Pt isotopes at least, ${ }^{196} \mathrm{Pt}$, more information soon became available when DiPrete et al. [DiP9:3] performed a low energy neutron scattering experiment with neutron beam energies between 2.0 and 3.5 MeV . DiPrete ct al. detected the $\gamma$-ray de-excitations from excited states and set constraints on the spin and parity assignments of the proposed octupole states based on $\gamma$-ray angular distributions. Twelve states were identified which could correspond to the same states populated by ( $p, p^{\prime}$ ). Unfortunately, Diprete ct al. could not determine exactly which states corresponded to the proposed octupole states. The difficulty in making a correspondence between the states seen by Diprete ct al. and the ones seen by Deason et al. stems from the non-selectivity of the neutron scattering reaction at low beam energies. Unlike protou scattering, which strougly populates single phonon states [Pet86], neutrou scattering populates all states at or below their excitation threshold energy in a statistical fashion to a near equal extent. Because there is no Coulomb barrier to overcome, states can be populated in ( $n, n^{\prime}$ ) which are only 100 or 200 keV above their excitation threshold, and spectral complications from higher-lying


Figure 4.4: Proton angular distributions compared to coupled channels calculations. The similarity between $3^{-}$and $4^{+}$states makes impossible an unambiguous spin assignment based on the proton angular distributions alone.
states are avoided. Thus, a number of states seen by Diprete et al. could correspond to the proposed octupole states.

The specific goal of the ${ }^{196} \mathrm{Pt}\left(n, n^{\prime} \gamma\right)$ experiment by Diprete ct al. was to identify the possible octupole states corresponding to those populated in ( $p, p^{\prime}$ ) by Deason et al. by extracting all possible spin and parity information from the measured properties of $\gamma$-ray decays. The minimal energy uncertainty provided by $\gamma$-ray detection permits much more reliable energy level determination than that obtainable from charged-particle detection such as ( $p, p^{\prime}$ ) [DeaSl] or ( $c, \epsilon^{\prime}$ ) [Pon92]. DiPrete et al. set specific criteria for a state seen in their ( $n, n^{\prime} \gamma$ ) reaction to be identified as an octupole state from the ( $\mu, p^{\prime}$ ) reaction. First, they required that the energy of the state fall within 15 keV of the energy reported by Deason ct al.. Deason et al. quote an energy uncertainty between $\mathrm{S}-10 \mathrm{keV}$ for their experiment. Secondly, DiPrete et al. required that the state have a spin and parity assignment of either $3^{-}$or $4^{+}$from their $\gamma$-ray decay properties. The states in question were known to have either $3^{-}$ or $4^{+}$assignments based on the $\left(p, p^{\prime}\right)$ angular distributions [Cot88a].

Of the 107 states observed in their ${ }^{1966} \mathrm{~F}^{\mathrm{t}}\left(n, n^{\prime} \gamma\right)$ reaction, Diprete et al. identified 12 states fitting these criteria for which a spin assignment of $3^{-}$or $4^{+}$was allowed by the $\gamma$-ray angular distributions, which could correspond to the possible octupole states seen in ( $p, p^{\prime}$ ). The candidate states proposed by Diprete et al. are shown in Figure 4.5 along with the states seen in $\left(p, p^{\prime}\right)$ to which they correspond.

### 4.2 Analysis of ${ }^{196} \operatorname{Pt}\left(p, p^{\prime} \gamma\right)$ Experiment

The intended goal of the present ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ coincidence experiment was to identify which four of the 12 candidate states observed in Reference [DiP93] were actually populated in the previous ( $p, p^{\prime}$ ) study of Deason et al. [Dea81]. Identifying these states allows energy level placements based on the highly reliable $\gamma$-ray decay energies as opposed to relying on tentative placements made from proton scattering

Possible Octupole States
Observed by ( $p, p^{\prime}$ )
(MeV)

$2.638 \longrightarrow$|  |
| :--- | :--- |
| 2.631 |
| 2.628 |



## Possible Corresponding States

 Observed by ( $n, n^{\prime} \mathrm{Y}$ )(MeV)

Figure 4.5: The proposed octupole candidate states in ${ }^{196} \mathrm{Pt} . ~ \gamma$-ray angular distributions allow for a $J^{\pi}=3^{-}$spin assignment for each state. Additionally, each state energy is, within error, consistent with one of the proposed octupole states seen in ( $p, p^{\prime}$ )
and electron scattering information. Although Ponomarev et al. have assigned three states in ${ }^{196} \mathrm{Pt}$ with $J^{\pi}=3-$ from an ${ }^{196} \mathrm{Pt}\left(\epsilon, c^{\prime}\right)$ study [Pon92], it is desirable to identify exactly which states their assignments refer to, using the best available information, namely, the $\gamma$-ray decay energies [DiP93].

It is possible that more than four of the candidate states were populated in the previous ( $p, p^{\prime}$ ) study. Since Deason et al. reported an energy resolution of no better than $2-8 \mathrm{keV}$, proton peaks interpreted as single peaks may, in fact, be multiple states which were unresolved in their experiment. However, there is no evidence in the present ( $\mu, p^{\prime} \gamma$ ) experiment for this occurring. It is further assumed that inelastic proton scattering of 12.7 MeV protons selects the same states seen in the $35 \mathrm{MeV}\left(p, p^{\prime}\right)$ reaction of Reference [DeaS1]. Table 4.2 summarizes the experimental findings of the present ${ }^{196} \mathrm{~F}^{\prime} \mathrm{t}\left(p, p^{\prime} \gamma\right)$ experiment. The table shows the populated states, $\gamma$-ray decays from these states, and measured branching ratios where possible. Figure 4.6 is an abbreviated energy level decay scheme showing the states populated in the present experiment. The following analysis will rely heavily on previous results, while taking advantage of the new coincidence information of the present experiment.

As an example of the quality of the protou- $\gamma$ coincidence information, Figure 4.7 shows $\gamma$-ray and proton projected spectra. The $\gamma$-ray spectrum is projected from a gate placed on proton events at 2245 keV and shows the $2245 \mathrm{keV} \gamma$-ray decay to the ground state, shown in the decay scheme of Figure 4.6. Likewise, the proton spectrum is projected from a gate placed on the $2.24 .5 \mathrm{keV} \gamma$-ray and shows the 2245 keV proton peak. The contaminant $1969 \mathrm{keV} \gamma$-ray is from the $\gamma$-ray decay chain of the 2262 keV state, which predominately decays to the 1969 keV state by a 293 $\mathrm{keV} \gamma$-ray. The 1969 keV state subsequeutly decays to the ground state by a 1969 $\mathrm{keV} \gamma$-ray. Due to our poor proton energy resolution ( 150 keV ), the 2262 keV state can not be separated from the 2245 keV state in the proton gating. The 2245 keV

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Figure 4.6: Abbreviated energy level decay scheme showing the states populated in the present ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ experiment.

Table 4.1: Euergy levels and $\gamma$-ray transitions observed in $E_{p}=12.7 \mathrm{MeV}$ ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ reaction.

| $\mathrm{E}_{\text {level }}(\mathrm{keV})$ | $\mathrm{E}_{\gamma}(\mathrm{keV})$ | Branching ratio | $\mathrm{E}_{f}(\mathrm{keV})$ | $\mathrm{I}_{i}^{\pi} \rightarrow I_{f}^{\pi}$ |
| ---: | ---: | :---: | :---: | :---: |
| $355.7(2)$ | $355.7(2)$ |  | 0 | $2^{+} \rightarrow 0^{+}$ |
| $688.7(2)$ | $332.9(2)$ |  | 356 | $2^{+} \rightarrow 2^{+}$ |
| $876.8(3)$ | $521.1(2)$ |  | 356 | $4^{+} \rightarrow 2^{+}$ |
| $1015.0(2)$ | $326.3(2)$ |  | 689 | $3^{+} \rightarrow 2^{+}$ |
| $1402.7(3)$ | $1047.0(3)$ |  | 356 | $0^{+} \rightarrow 2^{+}$ |
| $1447.0(2)$ | $758.4(2)$ | $12(4)$ | 689 | $3^{-} \rightarrow 2^{+}$ |
|  | $1091.1(3)$ | $88(17)$ | 356 | $3^{-} \rightarrow 2^{+}$ |
| $1535.8(3)$ | $847.1(3)$ | $17(9)$ | 689 | $4^{+} \rightarrow 2^{+}$ |
|  | $521.0(4)$ | $83(11)$ | 1015 | $4^{+} \rightarrow 3^{+}$ |
| $1677.2(5)$ | $1677.3(6)$ | $68(45)$ | 0 | $2^{+} \rightarrow 0^{+}$ |
|  | $1321.4(5)$ | $32(81)$ | 356 | $2^{+} \rightarrow 2^{+}$ |
| $1847.3(5)$ | $1491.6(5)$ |  | 356 | $2^{+} \rightarrow 2^{+}$ |
| $2047.0(4)$ | $1358.3(4)$ |  | 689 | $2^{+} \rightarrow 2^{+}$ |
| $2093.0(3)$ | $1737.3(3)$ |  | 356 | $\left(2^{+}\right) \rightarrow 2^{+}$ |
| $2245.6(7)$ | $2245.6(7)$ |  | 0 | $2^{+} \rightarrow 0^{+}$ |
| $2262.4(7)$ | $1246.7(7)$ | $67(15)$ | 1015 | $2^{+} \rightarrow 3^{+}$ |
|  | $1573.5(7)$ | $333(18)$ | 689 | $2^{+} \rightarrow 2^{+}$ |
| $2423.4(4)$ | $976.7(3)$ | $52(13)$ | 1447 | $3^{-} \rightarrow 3^{-}$ |
|  | $2067.6(7)$ | $48(22)$ | 356 | $3^{-} \rightarrow 2^{+}$ |

state is of some interest in a matter not related to octupole excitations, and will be briefly discussed in the following chapter.

Of the 4 higher-lying octupole states proposed by DiPrete et al. [DiP93], we see evidence for only one of the states. The evidence in support of an octupole inter-
pretation for this state, however, is very strong. Gating on inelastically scattered protons at 2423 keV reveals a $2067.6 \mathrm{keV} \gamma$-ray in coincidence with the protons. The $2067.6 \mathrm{keV} \gamma$-ray was previously reported by DiPrete et al. [DiP93] as decaying from a 2423 keV state to the $2_{1}^{+}$level (see Figure 4.6). A gate on the 2067.6 keV $\gamma$-ray shows this same coincidence as a proton peak at 2423 keV . Both $\gamma$-ray and proton projected spectra are shown in Figure 4.8 (a) and (b) respectively.

Figure 4.9 (a) also shows that, in additiou to the $2067.6 \mathrm{keV} \gamma$-ray decay, the 2423 keV state also decays by emission of a $976.7 \mathrm{keV} \gamma$-ray to the lowest-lying octupole state, the $3_{1}^{-}$state, seen in Figure 4.6. This decay has not been reported previously and lends certainty to the placement of the 2423 keV energy level. The proton projection showing the coincidence between the $976.7 \mathrm{keV} \gamma$-ray and the 2423 keV state is shown in Figure 4.9 (b).

Of the 12 candidate states identified in Reference [DiP93], the 2423 keV state is one of the three for which $J^{\pi}=4^{+}$is excluded by $\gamma$-ray angular distribution measurements. The protou angular distribution measurements of Reference [Cot88a] have already limited the spin assigument to either $3^{-}$or $4^{+}$. We can therefore unambiguously assigu $J^{\pi}=3^{-}$to the $242: 3 \mathrm{keV}$ level.


Figure 4.7: (a) $\gamma$-ray spectrum showing the $2245 \mathrm{keV} \gamma$-ray decay to the ground state and (b) the proton spectrum showing a peak from the 2245 keV state.


Figure 4.8: (a) $\gamma$-ray and (b) proton spectra showing the previously reported decay from the 2423 keV state by a $2067.6 \mathrm{keV} \gamma$-ray.


Figure 4.9: (a) $\gamma$-ray and (b) proton spectra showing the previously unreported decay from the 2423 keV state by a $976.7 \mathrm{keV} \gamma$-ray to the $3_{1}^{-}$state. The subsequent decays from the $3_{1}^{-}$state are also shown.

## CHAPTER 5

## DISCUSSION OF ${ }^{196} \mathbf{P t}\left(p, p^{\prime} \gamma\right)$ RESULTS

### 5.1 Octupole States in ${ }^{196} \mathbf{P t}$

Identification of an additional, higher-lying octupole state in ${ }^{196} \mathrm{Pt}$ lends considerable support to the argument made in Reference [Cot88b] for fragmentation of the low euergy octupole state. The present study shows that the low energy octupole strength is shared by ai least two states, at 1447 keV and 2423 keV . However, Cottle et al. proposed that three additional states ( 2608,2638 , and 2707 keV ) seen in the $35 \mathrm{MeV}{ }^{196} \mathrm{Pt}\left(p, p^{\prime}\right)$ experiment [Dea81] also share a significant amount of the octupole strength. Since the proton angular distributions for these other three states are nearly identical to that of the identified $2423 \mathrm{keV} 3_{2}^{-}$state, it is highly probable that the other three states also have $J^{\pi}=3^{-}$. We do not, however, have direct evidence for this from the present experiment.

There are several possible reasous why the other high-lying octupole states in ${ }^{196} \mathrm{Pt}$ at 2608.2638 , and 2707 keV proposed in Reference [Cot88b] were not observed in the present experiment. The assumption was made that the present experiment using 12.7 MeV protons would populate the same states seen in ( $p, p^{\prime}$ ) by Deason et al. using 35 MeV protous [Dea81]. It may be, however, that the incident proton beam energy of 12.7 MeV used in the present experiment is too low to excite the higher-lying octupole states. There were no clearly identifiable states, octupole or any other type, populated at higher energy in the present experiment than the $3_{2}^{-}$ octupole state which was identified. Secondly, two of the proposed octupole states (the 2608 keV and the 2707 keV states) have smaller B (E3) values (2.1 and $3.2 \mathrm{~W} . \mathrm{u}$.
respectively) than that for the 2423 keV state (4.4 W.u.) [Cot88b] and are therefore more difficult to observe. Lastly, it is possible that the 2423 keV state was more isolated in the proton spectrum than the other proposed states. The $\gamma$-ray spectrum from these states would not be as clean, making the high energy de-excitations more difficult to see than the weak 2067 keV decay that was observed. The resolution of the proton spectrum was not sufficient to determine whether this was true.

A detailed search for $\gamma$-rays from the other proposed octupole states reported by DiPrete et al. [DiP93] to the $2_{1}^{+}, 2_{2}^{+}, 3_{1}^{-}$, and $4_{1}^{+}$states was conducted by gating on the proton spectrum between 2.3 and 2.8 MeV and examining the resulting $\gamma$ ray spectra for the connecting transitions. Furthermore, $\gamma$-ray gates were set at the energies where the connecting transitions would occur, but no evidence of peak structure at the correct energies was seen in the resulting proton spectra.

### 5.2 Octupole States in Other Pt Isotopes

The observation of the 2423 keV state in ${ }^{196} \mathrm{Pt}$ is similar to a finding in a ${ }^{198} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ study of Yates et al. [Yat88]. They identified a state at 2603 keV which decayed to the $3_{1}^{-}$state at 1680 keV , and proposed, based on the occurrence of this transition, that the 2603 keV level has $J^{\pi}=3^{-}$. They associated this state with a 2611 keV state populated in the $35 \mathrm{MeV}{ }^{198} \mathrm{Pt}\left(p, p^{\prime}\right)$ reaction by Deason et al. [Dea81]. In a similar analysis to the one of ${ }^{196} \mathrm{Pt}$, Cottle et al. [Cot88b] proposed five higher-lying octupole states seen in the $35 \mathrm{MeV}{ }^{198} \mathrm{Pt}\left(p, p^{\prime}\right)$ reaction based on proton angular distribution measurements. These states, too, have angular distributious nearly identical to the identified $3.2 \cdot 2423 \mathrm{keV}$ state in ${ }^{196} \mathrm{Pt}$, and it is likely that they also have $J^{\pi}=3^{-}$. If this is the case, it provides evidence for octupole fragmentation in ${ }^{198} \mathrm{Pt}$ as well as ${ }^{196} \mathrm{Pt}$. Furthermore, Cottle et al. [Cot88b] proposed three higher-lying octupole states in ${ }^{194} \mathrm{Pt}$ having nearly identical proton angular distributions to the octupole states in ${ }^{196} \mathrm{P}$, and fragmentation of the low
energy octupole state may be present in ${ }^{194} \mathrm{Pt}$ as well. Direct experimental evidence is needed, however, to confirm these suppositions.

The similarity between the scattered proton angular distributions indicate that all of the proposed states in ${ }^{194,196,198} \mathrm{Pt}$ may have $J^{\pi}=3^{-}$. There remains, however, open discussion as to whether or not these states are collective octupole excitations. Yates et al. [Yat88] have suggested that the $3_{1}^{-}$states in ${ }^{194.196,198} \mathrm{Pt}$ are not vibrational octupole states, but are instead composed of two-quasiparticle states coupled to rotational excitations. However, reported $B\left(E 3 ; 0_{g s}^{+} \rightarrow 3_{1}^{-}\right)$values [Cot88a] derived from an analysis of the Deason ct al. ( $p, p^{\prime}$ ) data [Deasl] range from 6-9 W.u., indicating some degree of octupole collectivity for the states. The higher-lying $3^{-}$ states all have similar $B$ (E3) values, which yields a summed strength of $\approx 20 \mathrm{~W} . u$. for each of the Pt isotopes. The reported $\mathrm{B}(\mathrm{E} 3)$ values of Reference [Cot88a] are consistent with Coulomb excitation measurements for ${ }^{194} \mathrm{Pt}$ [Ron77].

Inelastic electron scattering measurements performed on ${ }^{196} \mathrm{Pt}$ by Ponomarev $e t$ al. [Pon92] also populate the three lowest $3^{-}$states in ${ }^{196} \mathrm{Pt}$ at $1447 \mathrm{keV}, 24: 31 \mathrm{keV}$, and 2638 keV . The authors assume collective excitations and assign $J^{\pi}=3^{-}$for the three states. The $B(E ; 3)$ values they extracted from scattered electron angular distributions are identical, within error, to the $B(E 3)$ values reported in Reference [Cot88a]. The values are given in Table 5.1. Although there may be some mixing from quasiparticle excitations, the $3^{-}$states clearly show octupole collectivity.

Identical $B(E 3)$ values from electron scattering, which probes the role of the protous, and from 35 MeV proton scattering, which primarily probes the neutron role, indicates that the amplitudes of the proton and neutron oscillations in the octupole vibrations are nearly equal. Identical proton and neutron oscillation amplitudes (that is, $\beta_{n}=\beta_{p}$ ) conforms to the simple isoscalar collective picture, and lends support to interpretation of these $3^{-}$states as collective excitations.

The $J^{\pi}=3^{-}$assignments of Ponomarev $c t$ al. [Pon92] provide a spin assignment of the proposed octupole states in ${ }^{196} \mathrm{Pt}$ which is independent of other measurements. However, from their scattered electron angular distributions, Ponomarev et al. cannot identify exactly which states they have populated, and rely, instead, on previously reported evergies from Deason it al. [Dea81]. We maintain that the superior energy resolution of $\gamma$-ray detection methods is the most reliable way of identifying these states, and that a coincidence experiment, such as ours, is required to make use of the existing $\gamma$-ray information to identify the precise energy of the $3^{-}$states in ${ }^{196} \mathrm{Pt}$.

Table 5.1: (Comparison of the transition strengths of the three lowest energy $3^{-}$ states in ${ }^{196} \mathrm{P}$ t from ( $p, p^{\prime}$ ) and ( $\epsilon, \epsilon^{\prime}$ )

| ${ }^{196} \mathrm{Pt}$ | 1447 keV | 2423 keV | 2638 keV |
| :---: | :---: | :---: | :---: |
| $\left(p, p^{\prime}\right)$ | $0.099(20) \epsilon^{2} b^{3}$ | $0.070(14) \epsilon^{2} b^{3}$ | $0.070(14) \epsilon^{2} b^{3}$ |
| $\left(\epsilon, \varepsilon^{\prime}\right)$ | $0.111(11) \epsilon^{2} b^{3}$ | $0.057(14) e^{2} b^{3}$ | $0.072(1: 3) \epsilon^{2} b^{3}$ |

### 5.3 Mixed Symmetry States in ${ }^{196} \mathbf{P} t$

The observation of a state at 2246 keV in ${ }^{196} \mathrm{Pt}$ in the present experiment raises a point which has little relation to octupole fragmentation in the Pt isotopes. It is, however, interesting in its own right, and deserves discussion here. A recent report by von Brentano et al. [vonBre96] indicates the first possible observation of a mixed symmetry "scissors mode" state in a $\gamma$-soft vucleus, ${ }^{196} \mathrm{Pt}$. The scissors mode is a vibrational mode in which the proton and neutron distributions oscillate with respect to oue another in a scissors type of motion. The mode, characterized by strong M1 transitions to the ground and $\varkappa_{1}^{+}$states, has been observed in a number of well-deformed and spherical uuclei, but never before in a $\gamma$-soft nucleus. Using
nuclear resonance fluorescence, von Brentano et ai. observed ten states between 2.2 and 3.5 MeV and suggested that these states may have $J^{\pi}=1^{+}$and are fragments of the scissors mode state. The observation of the 246 keV state in the present experiment indicates that the spin is more likely $J^{\pi}=2^{+}$.

Excitation by a photon type mechanism, such as nuclear resonance fluorescence. brings a very small amount of angular momentum to the reaction and thereby selects ouly $J=1,2$ states. On the basis of their measured branching ratios and excitation streugths, von Brentauo et al. report that a $\mathrm{l}^{+}$spin assignment is most favored for the 2246 keV state. The ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ reaction of the present experiment is also highly selective, however, with no known mechanism for strongly exciting a $1^{+}$state [Car85a, Dja85, Pet86].

The electromagnetic Ml operator has both spin and orbital components, which give rise to, respectively, the usual "spin" type excitation and a mixed symmetry orbital or "current" (the scissors mode) excitation. As noted in References [Car85a. Dja85], the cross section for exciting a purely orbital mixed symmetry $1^{+}$state with inelastic scattering of protons below 200 MeV is vanishingly small. In fact, as pointed out in Reference [Pet86], even the usual "spin" type $1^{+}$states are unlikely to be populated with 12.7 MeV protons. This appears to be true as no other $1^{+}$states were observed in the present experiment. Observation of the 2246 keV state in the present experiment, then, supports a $\cdot J^{\pi}=2^{+}$assignment. Assigning $J^{\pi}=2^{+}$to the 2246 keV state in no way contradicts the assertion of von Brentano ct al. for their observation of the first scissors mode state in a $\gamma$-soft nucleus; it simply indicates that the 2246 keV state is not one of the fragments of the $1^{+}$scissors mode state.

There exists the possibility that the 2246 keV state may still be a mixed symmetry excitation. Iachello [Iac84] predicted that the lowest mixed symmetry state in an $O(6)$ nucleus, of which ${ }^{196} \mathrm{Pt}$ is the best known example, would have $J^{\pi}=2^{+}$and a large E2 transition to the gromed state. This is consistent with what was seen in the
present experiment. With the exception of the $2_{1}^{+}$state, the fully symmetric states in an $O(6)$ nucleus are not predicted by Iachello to have significant E2 transitions to the ground state. The E 2 transition strength for the 2246 keV state decaying to the ground state is 2.6 W.u. (from width measured by von Brentano et al. [vonBre96]), making this state a likely candidate for the $2^{+}$mixed symmetry excitation.

## CHAPTER 6

## THEORETICAL OVERVIEW FOR $0_{g s}^{+} \rightarrow 2_{1}^{+}$TRANSITION IN ${ }^{20} \mathbf{O}$

As was discussed in Chapter 2, the bulk of nuclear behavior exhibits properties of both the idealized pictures of the shell model and the collective model. While the data on $2_{1}^{+}$states in some of the nuclei of the $12 \leq A \leq 26$ mass region suggest collective interpretations, our conclusions indicate that the shell structure has a strong effect on the behavior of the $0_{g s}^{+} \rightarrow 2_{1}^{+}$transition.

### 6.1 Structure of First Excited States in the $12 \leq A \leq 26$ Mass Region

The two most common descriptions of macroscopic collective behavior are as rotational and vibrational excitations. Collective $2_{1}^{+}$states, the primary concern of this chapter, can be easily described by both types. Rotational and vibrational behavior is usually distinguished experimentally by the systematic spacing of energy levels corresponding to successive quanta of vibrational or rotational excitation. Figure 6.1 shows typical energy level spacings for vibrational and rotational behavior. While some nuclei in the $12 \leq A \leq 26$ mass region clearly exhibit rotational behavior, such as ${ }^{20,22} \mathrm{Ne}$ and ${ }^{24,26} \mathrm{Mg}$, the known energy levels of ${ }^{20} \mathrm{O}$, shown in Figure 6.2, suggest vibrational behavior in this nucleus. In particular, the energies of the $2_{1}^{+}$and $4_{1}^{+}$states in ${ }^{20} \mathrm{O}$ are very near the harmonic vibrational limit, with $E\left(4_{1}^{+}\right) / E\left(2_{1}^{+}\right)=2.18$. Elliott et al. have pointed out that the energy spectrum of ${ }^{20} \mathrm{O}$ indicates vibrational behavior, while that of the neighboring nucleus, ${ }^{20} \mathrm{Ne}$, is interpreted to be rotational in their IBM treatment of $A=20$ uuclei [Ell80]. Because of the nearly equal energy level spacing, and since stable nuclear deformation
generally only occurs in the ground state when a nucleus has both protons and neutrons in the valence shells [Cas81], we choose to model the $2_{1}^{+}$state of the singly closed shell nucleus ${ }^{20} \mathrm{O}$ as a quadrupole vibrational state.

Vibrational excitations occur in energy quanta of a given multipole order known as vibrational phonous. The multipole order also determines the angular momentum quanta carried by the vibrational phonon. In the case of a quadrupole vibration, the angular momentum of the excited phonon is $2 \hbar$. Phonons of the same or different multipolarities can couple to each other, producing multiplets of excited states. An example of this is shown in Figure 6.1, where quadrupole phonons are coupled to produce multiplets of degenerate positive parity states. In physical cases, the degeneracy of the multiplets is broken because of residual interactions between the phonons. Phonon-phonon interactions also account for the deviations from a perfectly harmonic level spacing.

From a microscopic point of view, vibrational excitations are coherent superpositions of quasiparticle excitations. However, in the $A<40$ mass region, there are relatively few single-particle configurations available with which to build vibrational excitation modes. In spite of this, many of these nuclei still exhibit the enhanced transition probabilities characteristic of collective behavior. Due to the low binding energy of these uuclei, $\left({ }^{20} \mathrm{O}\right.$, for example, has a neutrou separation energy of 7.6 MeV ) the number of excited states is oftentimes not sufficient to clearly determine whether the collectivity is vibrational or rotational based on energy level spacing (see Figure 6.2). The coupled channels approach we have used to analyze the scattered proton angular distributions for the ground state and $2_{1}^{+}$state in ${ }^{20} \mathrm{O}$ is not very sensitive to the choice of a vibrational or rotational model. The coupled channels calculations were performed for both models for ${ }^{20} \mathrm{Ne}$, with very little difference to the results whether the quadrupole deformation was treated as dyuamic or static.


Figure 6.1: Typical energy level spacings from vibrational and rotational collective behavior

We have therefore chosen, in the present analysis, to model the $2_{1}^{+}$state in ${ }^{20} \mathrm{O}$ as a vibrational collective excitation, following the interpretation of Elliott et al. [Ell80].

Although there is no attempt in the present work to describe the nuclear behavior microscopically in an explicit way, the shell model is very useful in understanding collective excitations in the $12 \leq A \leq 26$ mass region. Shell model descriptions typically treat closed shells as inert cores which contribute little to the excitation. and only the valence nucleons outside of closed shells are free to populate higher orbitals. To quantify the isospin behavior, the isospin quantum number $T$ and its projection onto au arbitrary isospin symmetry axis $T_{z}$ are used. The ground state isospin for even-even uuclei is the minimum value allowed for a given $T_{z}=1 / 2(Z-N)$


Figure 6.2: Known energy levels in ${ }^{20} \mathrm{O}$.
projection. Any transition with $\Delta T=0$ is purely isoscalar, and any with $\Delta T=1$ is purely isovector. Ofteutimes, though, a trausition will have both isoscalar and isovector compouents.

The choice of a particular nuclear model can make a huge difference in predicting relative neutron and proton contributious in these transitions. In the simplest shell model picture, vibrations cousist solely of valeuce uucleon excitations. Vibrations in a nucleus with a closed proton shell, like ${ }^{20} \mathrm{O}$ for example, are viewed as entirely due to the valence neutrons. The ratio of neutron to proton transitiou matrix elements $M_{n} / M_{p}$ for this transition would be infinite in this example, and the vibration would be exclusively from neutron excitations. The isoscalar and isovector transition matrix elements, $M_{0}$ and $M_{1}$ respectively, are related to the proton and neutron transition matrix elements by

$$
\begin{align*}
& M_{0}=M_{p}+M_{n} \\
& M_{1}=M_{p}-M_{n} \tag{6.1}
\end{align*}
$$

although other sign conventions are used [Ber79, Bro82, Pet93]. In the simplest collective picture, however, the uucleus is thought of as a homogeneous neutronproton fluid with the protons and neutrons oscillating with equal amplitudes. In the example of ${ }^{20} \mathrm{O}$ just used, the same transition which gave $M_{n} / M_{p}=\infty$ in the simplest shell model, would give $M_{n} / M_{p}=N / Z$ in the simplest collective model. The participation of excitations of the core nucleons, known as core polarization, resolves to some extent, the drastic discrepancy between these two pictures [Bro7.5, Bro80, Bro82].

In a more realistic shell model, the closed shell core is not entirely inert, and the participation of core uucleou excitations serves to reduce $M_{n} / M_{p}$ from infinity in singly closed proton shell uuclei and increase $M_{n} / M_{p}$ from zero in singly closed neutron shell nuclei for the $0_{g s} \rightarrow 2_{1}^{+}$transition [Ber83]. The magnitude of the core
polarization is always greater for the core nucleon type which is unlike the valence nucleon type because the residual interaction between unlike nucleons is stronger than that between like uucleous [Ber81, Ber83]. Although low-lying collective $2_{1}^{+}$ states are primarily composed of $0 \hbar \omega$ excitations from the valence nucleons, higherlying $1 \hbar \omega$ and $2 \hbar \omega$ excitations of the core contribute to the excitation [Ber83]. The degree of mixing between core nucleon excitations and valence nucleon excitations is quantified in shell model treatments by introducing core polarization parameters which connect the proton and neutron transition matrix elements $A_{p}$ and $A_{n}$ in the shell model valence space with the full transition matrix elements $M_{p}$ and $M_{n} . M_{p}$ and $M_{n}$ can be defined as

$$
\begin{align*}
& M_{p}=A_{p}\left(1+\epsilon_{p p}\right)+A_{n} \epsilon_{p n} \\
& M_{n}=A_{n}\left(1+\epsilon_{n n}\right)+A_{p} \epsilon_{n p} \tag{6.2}
\end{align*}
$$

where $\epsilon_{a b}$ signify the strength of the coupling between valence nucleons $b$ and the core nucleous a [Bros2].

A systematic coupling scheme has been used by Bernstein et al. [Ber81] to set the core polarization parameters so that the experimental energies of the isoscalar and isovector giant quadrupole resonances are reproduced. Their scheme reproduces extracted ratios of transition matrix elements, $M_{n} / M_{p}$, for $0_{g s}^{+} \rightarrow 2^{+}$transitions in singly closed shell nuclei between $18 \leq A \leq 208$. The same scheme has been extended by Madsden and Brown [Mad84] to include open shell nuclei between $N=$ 50 and $N=82$. Brown et al. [Bro82] have also performed shell model calculations for $2_{1}^{+}$states in selected $8 \leq A \leq 20$ uuclei, although only a few calculations exist for nuclei away from stability. Their calculations for nuclei close to stability assume an inert ${ }^{16} \mathrm{O}$ core with $n=A-16$ valence nucleons in a complete $0 d_{5 / 2}, 1 s_{1 / 2}, 0 d_{3 / 2}$ basis of Chung and Wildenthal [Bro80b], notated as $(s d)^{n}$. The results from the calculations of Brown et al. [Bro82] will be discussed in Chapter 9. Brown et al.
[Bro82] concede that their calculations are somewhat model dependent, and they point specifically to their assumption of harmonic oscillator radial wave functions.

### 6.2 Probe Sensitivity

To investigate differences between proton and neutron contributions it is often necessary to use two independent experimeutal probes, one sensitive to neutrons and one sensitive to protons. With the exceptiou of electromagnetic (EM) probes, which interact with the protons only, every probe has its own unique interaction with both protous and neutrous. Thus, different probes have a variety of sensitivities to the neutron contribution relative to the proton contribution of the $0_{g s}^{+} \rightarrow \mathscr{2}_{1}^{+}$transition. Each probe, therefore, measures a mixture of the transition matrix elements $M_{n}$ and $M_{p}$ [Ber81]. If the relative sensitivity of the probe to neutron and proton contributions is known, however, the transition operator $O_{\lambda \mu}^{F}$ can be decomposed into neutron and proton operators such as

$$
\begin{equation*}
O_{\lambda \mu}^{F}=b_{n}^{F} O_{\lambda \mu}^{F n}+b_{p}^{F} O_{\lambda \mu}^{F p}, \tag{6.3}
\end{equation*}
$$

where $b_{n}^{F}$ and $b_{p}^{F}$ are the neutron and proton external field strengths for a given probe $F$. The multipole matrix element measured by a given probe is then

$$
\begin{align*}
M_{F} & =\left\langle\nu_{f} J_{f} T_{Z} A\left\|O_{\lambda}^{F}\right\| \nu_{i} J_{i} T_{Z} A\right\rangle \\
& =b_{n}^{F} M_{n}+b_{p}^{F} M_{p} . \tag{6.4}
\end{align*}
$$

For two different probes $F$ and $G$ having different neutron and proton sensitivities, $M_{n}$ and $M_{p}$ can be extracted from $M_{F}$ and $M_{G}$.

In the case of scattering probes, with which the present work is concerned, the extraction is typically done by relating the transition matrix element of a given probe to the RMS multipole deformation length $\delta_{\lambda}^{F}=r_{\lambda}^{F} \beta_{\lambda}^{F}$. $\delta_{\lambda}^{F}$ is experimentally obtainable by a coupled channels or folding model analysis of the elastic and inelastic
differential cross sections. The mean radius $r_{\lambda}^{F}$ depends on the particular probe, and the $\beta_{\lambda}^{F}$ are RMS deformation parameters which correspond to the multipole order expansiou of the nuclear surface R in Equation 1.1. Just as the matrix element $M_{F}$ for a given probe is decomposed into ueutrou and proton components, so is the RMS deformation length as

$$
\begin{equation*}
\delta_{\lambda}^{F}=\frac{b_{n}^{p} N \delta_{\lambda}^{n}+b_{p}^{p} Z \delta_{\lambda}^{p}}{b_{n}^{p} N+b_{p}^{p} Z} . \tag{6.5}
\end{equation*}
$$

The ratio of neutron to proton transition matrix elements as measured by two different probes $F$ and $G$ is then

$$
\begin{equation*}
\frac{M_{n}}{M_{p}}=\frac{N \delta_{\lambda}^{n}}{Z \delta_{\lambda}^{p}} \tag{6.6}
\end{equation*}
$$

Bernstein et al. [Ber81] have compiled approximate sensitivity ratios $b_{n}^{F} / b_{p}^{F}$ for various experimental probes in specific energy ranges which allow us to use the explicit form of Equation 6.5 to obtain
where

$$
\begin{equation*}
\epsilon_{G i}^{F}=\frac{1+\frac{N b^{F}}{Z b_{p}^{F}}}{1+\frac{N b^{G} F}{Z b_{p}^{F}}} . \tag{6.8}
\end{equation*}
$$

$M_{n} / M_{p}$ values which incorporate these sensitivity parameters are consistent among the experimental probes to within $15 \%$ [Ber8I] over a mass range of $18 \leq A \leq 208$. The $b_{n}^{F} / b_{p}^{F}$ ratios listed by Bernstein et al. are given in Table 6.1.

For the present work, which focuses on inelastic proton scattering, the ratio $b_{n}^{N} / b_{p}^{N}$ is derived from phenomenological optical model potentials for elastic nucleon scattering. The treatment, taken from Satchler [Sat83], is general for any type of nucleon scattering in a $10-50 \mathrm{MeV}$ euergy range. A general optical potential $V_{N}$ for elastic nucleon scattering can be given by the Lane potential [Sat64, Mad75], and written in terms of isoscalar and isovector componeuts as

$$
\begin{equation*}
V_{N}=V_{0}+V_{1} \frac{4}{A} \vec{\tau} \cdot \vec{T} \tag{6.9}
\end{equation*}
$$

Table 6.1: Ratios of neutron to protou sensitivity of various probes. [Ber81]

| Probe (F) | Euergy [MeV] | $b_{n}^{F} / b_{p}^{F}$ |
| :--- | :---: | ---: |
| $E M$ | - | 0 |
| $\left(p, p^{\prime}\right)$ | $10-50$ | $\approx 3$ |
| $\left(n, n^{\prime}\right)$ | $10-50$ | $\approx \frac{1}{3}$ |
| $\left(\pi^{+}, \pi^{+^{\prime}}\right)$ | $160-200$ | $\approx \frac{1}{3}$ |
| $\left(\pi^{-}, \pi^{-\prime}\right)$ | $160-200$ | $\approx 3$ |
| $\left(p, p^{\prime}\right)$ | 800 | 0.83 |
| $\left(p, p^{\prime}\right)$ | 1000 | 0.95 |
| $\left(\alpha, \alpha^{\prime}\right)$ | All | 1 |

with $\vec{\tau}$ and $\vec{T}$ as the respective isospin operators for the incident nucleon and target nucleus, and $V_{0}$ and $V_{1}$ respectively containing the isoscalar and isovector potential form factors and well depths. Using the isospin projections, $\tau_{z}=\frac{1}{2}$ for protons and $\tau_{z}=-\frac{1}{2}$ for neutrous, we can specify the nucleon scattering potential $V_{N}$ for neutron and proton scattering as

$$
\begin{align*}
& V_{p}=V_{0}+V_{1} \frac{(N-Z)}{A}  \tag{6.10}\\
& V_{n}=V_{0}-V_{1} \frac{(N-Z)}{A} . \tag{6.11}
\end{align*}
$$

The potential simplifies with the assumption of identical form factors for the isoscalar and isovector componeuts, and the assumption that the ratios of real to imaginary parts of $V_{0}$ and $V_{1}$ are identical. An additional assumption must be made about the ratio of isoscalar to isovector potential strength $V_{0} / V_{1}$. The existence of the deuteron, and the absence of a bound neutron-neutron or proton-proton state, indicates that the isoscalar potential is considerably stronger than the isovector. A lower limit of $V_{0} / V_{1}=-3$ is made by neglecting odd relative angular momentum interactions [Car85b]. However, more realistic potentials give the value $V_{0} / V_{1}=-2$
[Car85b, Ber81, Mad75]. Using the more realistic estimate of $V_{0} / V_{1}=-2$ gives the probe sensitivity to protous and neutrons as the coefficients of $N$ and $Z$ from proton scattering,

$$
\begin{align*}
V_{p} & \approx V_{0}\left[1+\frac{(N-Z)}{2 A}\right]  \tag{6.12}\\
& \approx \frac{V_{0}}{2 A}(3 N+Z),
\end{align*}
$$

and from neutron scattering,

$$
\begin{align*}
V_{n} & \approx V_{0}\left[1-\frac{(N-Z)}{2 A}\right]  \tag{6.1:3}\\
& \approx \frac{V_{0}}{2 A}(N+3 Z)
\end{align*}
$$

The $N$ and $Z$ coefficients show that proton scattering is a factor of three more sensitive to the neutron distribution ( $b_{n}^{\left(p, p^{\prime}\right)}=3, b_{p}^{\left(p, p^{\prime}\right)}=1$ ), and that neutron scattering is a factor of three more seusitive to the proton distribution $\left(b_{n}^{\left(n, n^{\prime}\right)}=1\right.$. $b_{p}^{\left(n, n^{\prime}\right)}=3$ ). Ratios of $b_{n}^{F} / b_{p}^{F}$ are givell in Table 6.1.

Table 6.1 also lists probe sensitivities for pion scattering $\left(\pi^{+}, \pi^{-}\right)$between $160-$ 200 MeV , an independent method of extracting the $M_{n} / M_{p}$. The pion scattering method is used in the present work as a comparison of the experimental method and is discussed in Chapter 9.

### 6.3 Coupled Channels Calculations

As was mentioned in the preceding sections, the dynamic, RMS quadrupole deformation parameter $\beta_{2}$ is usually extracted from a measured angular distribution by means of a coupled channels or folding model analysis. We have chosen, for the $0_{g . s}^{+} \rightarrow 2_{1}^{+}$transition in ${ }^{20} \mathrm{O}$, to extract $\beta_{2}$ from the elastic and inelastic proton scattering angular distributions by way of a coupled channels calculation. The computer code CHUCK [Kunz] was used for this purpose. CHUCK performs a
numerical integration to simultaneously solve a set of $n$ coupled equations of the form

$$
\begin{equation*}
\left[\frac{d^{2}}{d r_{c}^{2}}-\frac{l_{c}\left(l_{c}+1\right)}{r_{c}^{2}}+\left(k_{c}^{2}-U_{c c}\right)\right] \chi_{j_{c} l_{c}}^{J}\left(k_{c} ; r_{c}\right)=\sum_{c^{\prime} \neq c} \frac{2 \mu_{c}}{\hbar^{2}} U_{c c^{\prime}}^{\prime} \chi_{j_{c^{\prime}} l_{c^{\prime}}}^{J}\left(k_{c^{\prime}} ; r_{c^{\prime}}\right), \tag{6.14}
\end{equation*}
$$

where $\chi_{j_{c} l_{c}}^{J}\left(k_{c} ; r_{c}\right)$ are the available channel wavefunctions describing the relative motion of the projectile and the target with reduced mass $\mu_{c}$ for a given channel $c$ by the relative coordinates $\vec{r}_{c}$ and wavenumber $k_{c}$. The diagonal and off-diagonal terms of the optical potential are denoted by $U_{c c}$ and $U_{c c^{\prime}}^{\prime}$, respectively. The channel wavefunctions are found by a partial wave expansion of the allowed orbital angular momenta $l_{c}$ of the projectile about the target. The total angular momentum $J$ is the sum of the orbital angular momentum, $l_{c}$, and the intrinsic moments of inertia $s_{c}$ and $I_{c}$ of the projectile and the target respectively, with the coupling scheme $\vec{j}_{c}=\vec{l}_{c}+\vec{s}_{c}$ for every allowed channel $c . \quad \gamma_{c}$ is the reduced mass of the projectile and target system. The full potential consists of both diagonal terms, $U_{\text {cc }}$, and off diagonal terms, $U_{c c^{\prime}}^{\prime}$. $U_{c c}$ is a many-component optical potential based on an optical model analysis of elastic cross sections, while $U_{c c^{\prime}}^{\prime}$ is a coupling term describing the effective interaction for transitions between channels $c$ and $c^{\prime}$. The code CHUCK allows for user-defined nuclear poteutials, as well as the Coloumb interaction for both $U_{c c}$ and $U_{c c^{\prime}}^{\prime}$. The details of the potential forms used in the present analysis, as well as the potential well depths, is given in Chapter 8.

CHUCK produces a physically meaningful solution with the application of two boundary conditions, one at the origin of the projectile and target system, and one at the far field range of the Coloumb interaction. In general, the solution is a linear combination of the independent $n$ solutions of the coupled equations. The output from CHUCK can be directed to give an angular distribution for all included channels in the center of mass of the projectile and target system, providing a comparison to the observed experimental augular distribution. In the present
analysis, $\beta_{2}$ is used as a user-defined coupling parameter which best reproduces the experimentally observed angular distribution.

## CHAPTER 7

## ANGULAR DISTRIBUTION MEASUREMENT FOR ${ }^{20} \mathbf{O}$

An experiment to measure the angular distribution of protons scattered from the radioactive beam ${ }^{20} \mathrm{O}$ in inverse kinematics was performed at the National Superconducting Cyclotron Laboratory of Michigan State University. Protons from 1.6 $\mathrm{mg} / \mathrm{cm}^{2}$ and $3.3 \mathrm{mg} / \mathrm{cm}^{2}$ polypropylene targets were scattered from a $30 \mathrm{MeV} / \mathrm{A}$ beam of ${ }^{20} \mathrm{O}$. The polypropylene targets provide a source of protons from which the ${ }^{20} \mathrm{O}$ beam is scattered. The inverse kinematics experiment is equivalent to performing 30 MeV proton scattering from a target of ${ }^{20} \mathrm{O}$.

### 7.1 Beam Production

The present $p\left({ }^{20} \mathrm{O}, p\right)^{20} \mathrm{O}^{*}$ experiment was performed at the National Superconducting Cyclotron Laboratory, one of the few facilities equipped to deliver high intensity, high energy radioactive beams. A layout of the laboratory, including the K1200 cyclotron, the Al200 mass fragment separator, and the $S 2$ experimental vault, is shown in Figure 7.1. Radioactive beams at the NSCL are produced by means of particle fragmentation.

A primary stable beam of ${ }^{22} \mathrm{Ne}$ was produced in an electron cyclotron resonance (ECR) ion source and injected into the K1200 cyclotron. There the primary beam was accelerated to an energy of $65 \mathrm{MeV} / \mathrm{A}$ and steered into the entrance of the Al200 mass fragment separator [She91]. The primary beam collided with a 360 $\mathrm{mg} / \mathrm{cm}^{2}$ water-cooled ${ }^{9}$ Be target positioned at the entrance of the Al200, shown in Figure 7.2.


Figure 7.1: The layout of the NSC'L experimental facility. The figure shows the K1200 cyclotron where the primary beam is accelerated, the A1200 mass fragment separator for production and selection of the radioactive secondary beam, and the S2 experimental vault at the tail end of the RPMS fragment separator where the $p\left({ }^{20} \mathrm{O}, p\right)^{20} \mathrm{O}$ scattering experiment was performed.

The primary beam fragmented upon impact with the ${ }^{9}$ Be target into a near continuum of both stable and unstable nuclei. The nucleus used as a secondary beam, in this case the radioactive beam ${ }^{20} \mathrm{O}$, was selected in the Al200 by time of flight and energy loss characteristics. Degrading wedges placed at the dispersive image of the Al200 allowed isotope separation. For the present experiment, a $1 \%$ momentum acceptance was allowed by the placement of a $1.9 \mathrm{mrad}{ }^{12} \mathrm{C}$ degrading wedge at the second dispersive image pusition. Owing to the high initial beam energy, no re-acceleration of the secondary beam was required. The secondary beam


Figure 7.2: Al200 mass fragment separator at the NSCL. The Al200 is used to identify and select secondary beams after fragmeutation of the primary beam.
${ }^{20} \mathrm{O}$ was produced at $30 \mathrm{MeV} / \mathrm{A}$. The beam was nearly $99 \%$ pure, and yielded roughly 30,000 particles/second (pps) on target.

### 7.2 Experimental Details

### 7.2.1 Bearn Tracking and Identification

The secondary beam was tracked by two position sensitive parallel-plate avalanche counters (PPACS) [Swa94] placed at 1 m and 2 m upstream of the target position. The PPAC positions, along with a schematic of the other detector positions, can be seen in Figure 7.3. The PPACS provide accurate position sensitivity in two dimensions over a $10 \times 10 \mathrm{~cm}^{2}$ active area. Low pressure iso-octane gas constrained between two sheets of mylar provided an ionization read out by $0.4 \mathrm{mg} / \mathrm{cm}^{2}$ aluminized polyester and polypropylene electrodes affixed to each mylar sheet. Two independent read outs were made for each of the two position dimensions so that a simple charge division determined the position. Fast preamplifiers capable of count


Figure 7.3: Schematic of the experimental arrangement.
rates up to $5 \times 10^{4}$ events/second were used. Initial position calibration measurements were taken with an $\alpha$ source and a mask placed over the PPAC front surface. The mask was a square grid of 2 mm diameter holes spaced 1 cm apart. A position resolution of 1 mm was observed from this calibration. A subsequent in-beam calibration was performed which confirmed the resolution and alignment of the PPACs on the optical beam axis.

The beam was stopped by a fast/slow plastic phoswich placed at a zero degree forward scattering angle 1 m downstream of the target position. The zero degree detector had an angular acceptance of 2.5 degrees which was sufficient to detect the
entire angular scattering of the ${ }^{20} \mathrm{O}$ beam. $\Delta E-E$ energy loss measurements made from comparing the fast and slow zero degree detector signals allowed a moderate degree of beam identification based on mass and charge number. The high purity ( $99 \%$ ) of the secondary beam, however, made this feature somewhat unnecessary. For further identification, the zero degree detector was also used to record the time of flight (TOF) of beam particles. A PIN diode located at the Al:200 exit provided the start of the time of flight measurement. Count rates up to $200,000 \mathrm{pps}$ were possible with the zero degree detector. Sample $\Delta E-E$ and TOF plots showing the high beam purity are given in Figure i.4.

### 7.2.2 FSU-MSU Charged-Particle Telescope Array

A charged particle telescope array consisting of 8 three stage Si strip-Si PIN-CsI particle telescopes was used to detect scattered protons. Each stage of the telescopes has a $5 \times 5 \mathrm{~cm}^{2}$ active area. The first stage of the telescope is a $300 \mu \mathrm{~m}$ thick position sensitive Si strip detector. The detector face is segmented into 16 individual strips, 3 mm in width, which are read out individually. A $500 \mu \mathrm{~m}$ thick PIN diode makes up the secoud telescope stage, and a 1 cm thick ('si scintillator composes the last stage. Four separate photodiodes attached to the CsI crystal by a clear RTV epoxy were used to collect light from the scintillator. The multi-stage telescope provides not ouly a means for angular distribution measurements, which could just as easily be done with a single--layer position sensitive detector, but also a means for particle identification. Particles fully stopped in the Si strip detector were identified by time of flight. Higher energy particles which penetrate into two or more layers of the detector were identified by $\Delta E-E$ measurements between the different telescope layers. The various means of particle identification allows filtering of deuterons and alpha particles from the collected data.

Preamplifiers for the PIN diodes and CsI detectors were placed inside the scattering chamber to reduce noise in the signals. Due to the many distinct channels


Figure 7.4: Sample $\Delta E-E$ and TOF zero degree detector spectra showing the high purity of the beam. The top plot is the energy loss in the fast plastic versus the energy in the slow plastic. Only a single isotope $\left({ }^{20} \mathrm{O}\right)$ can be seen. The bottom plot is the beam TOF between the Al200 exit and the zero degree detector. The spreading of the characteristic energy loss point is caused by an unstable power supply producing fluctuations in the detector bias voltage
required by the Si strip detectors, it was not practical to house the Si strip preamplifiers inside the scattering chamber. Those preamplifiers were instead located just outside of the chamber. The placement of the Si strip detector preamplifiers outside the chamber led to a large increase in noise to the signals, requiring noise threshold settings close to 1 MeV . The PIN diode thresholds, on the other hand, were around 250 keV . All of the preamplifiers used with the charged-particle telescope array were built at the NSCL.

The physical positioning of the telescopes was done so that the Si strips lie tangent to circles of constant scattering angle. The scattering angle is measured in the laboratory frame with respect to a coordinate system having its origin at the target and $z$-axis coinciding with the optical beamline axis. Each telescope is mounted on one of two vertical bars by an aluminum wedge which aligus each strip along a constant scattering angle. Figure 7.5 (as well as Figure 7.3 ) illustrates the telescope array positioning. The telescopes were mounted 28 cm from the target position in the present experiment. At this position, each telescope could detect an angular range of $10^{\circ}$, with each individual strip seeing a range of $0.6^{\circ}$ in the laboratory frame. Due to the strips not being curved to match the constant scattering angular curves, and due to a large area of the target being illuminated by the beam, each strip actually detected a rauge of about $0.85^{\circ}$, thereby introducing a modest amount of uncertainty in the overall augular resolution.

Two major concerns were taken into account in choosing the scattering angles at which to place the telescopes. First, the angular range must cover a region where both elastically and inelastically scattered protons are expected from a kinematics calculation, and where they are well separated in energy. Secondly, the angles chosen must be in a region for which the scattering cross section is reasonably high. Additionally, the array as a whole must cover a large enough angular range to yield a multi-point angular distribution. To meet these criteria, three of the telescopes


Figure 7.5: Positioning of the FSU-MSU charged particle telescope array. Each telescope is a three stage detector consisting of a position sensitive Si strip detector, a PIN diode, and a CsI scintillator. Telescopes 1,2 and 3, on the left side, are centered at $75^{\circ}$ with respect to the beam axis, and telescope 4 is at $60.5^{\circ}$. On the right, telescopes 5,6 and 7 are centered at $70^{\circ}$, and telescope 8 is at $84.5^{\circ}$.
were centered at $70^{\circ}$ on one side of the target, three more were centered at $75^{\circ}$ on the other side, and the remaining two telescopes were centered at $84.5^{\circ}$ and $60.5^{\circ}$. Even without the use of the detectors at $84.5^{\circ}$ and $60.5^{\circ}$, which proved unreliable due to electronics difficulties, measurement over a range of $65^{\circ}-82^{\circ}$ in the laboratory frame was possible. This corresponds to an angular range of roughly $20^{\circ}-45^{\circ}$ in the center of mass frame.

### 7.3 Electronics

A diagram showing the electronics used for beam tracking and identification is shown in Figure 7.6. Each of the two PPAC detectors provides four signals corresponding to left, right, up, and down for determining the position of a passing beam particle. Positions are determined by a charge ratio between the respective left and right siguals and up and down signals. The signals are first attenuated and then shaped in a slow amplifier before being sent directly to an ADC. A fast timing signal is also produced in a separate amplifier for any PPAC event. The timing signal is discriminated by a (CFD, delayed by 300 ns , and read by a time to digital converter (TDC).

The zero degree detector signal is decomposed into two components, one each for the fast and slow euergy signals, and are read directly by FERAs after the preamplifier. A fast timing signal is also produced in the zero degree detector preamplifier which is used in the master logic signal, described below. The timing signal is discriminated to filter noise events. Three identical outputs are taken from the discriminator. One is used in the master logic circuit. The other two are sent to scaler counters, one directly and the other after being rate divided by a factor of 1,000 , for absolute cross section measurements. Timing signals were also taken from the beamline PIN diode at the exit of the K1200 and from the RF cyclotron
frequency. Both of these timing siguals, and backup timing signals for each, are read by TDCs.

The electronics for the particle telescopes are considerably more extensive than the beam tracking electronics, mostly due to the number of independent channels required. Figure 7.7 is a schematic diagram of the electronics used for a single particle telescope, including Si strip, PIN diode, and CsI signals. Energy signals from each of the 16 Si strips are first amplified in a 16 channel preamplifier, then shaped in a slow amplifier before being read by a peak sensing ADC. The timing signals, also generated in the 16 channel preamplifier, are shaped by an inverting fast amplifier and discriminated by a ('FD). Two outputs from the CFD are delayed by 150 us and used for measuring time signals, one via a common start TF( and read by a FERA for TOF measurements, and the other sent to a scaler counter. A third CFD output is used in the master logic through a logical OR condition with the other telescope strip timing signals. The PIN diode and C'sI siguals are handled in a similar fashion. The PIN sigual is preamplified, then shaped in a combined slow/fast amplifier which gives both energy and timing signals. The energy signal is read directly by an ADC, while the timing sigual is discriminated, delayed by 200 ns, and read by a TDC. The Csl signals are processed identically, with the exception that the timing signal is read by a scaler counter.

The master logic circuit is in direct communication with the data acquisition hardware and software, and ultimately controls which events are written to tape. When the master logic coincidence requirement is met, which will be discussed immediately following, all of the aforementioned signals are read out together. The coincidence conditiou for a "good" event requires detection of a particle in any of the Si strips and detection of a beam particle in the zero degree detector. A coincidence logic unit (CLU), shown in Figure 7.8 along with the rest of the master coincidence electronics, receives delayed gates from the OR of all strip detectors and the zero
degree detector. If both signals are preseut, the Master CLU outputs a positive coincidence sigual which is passed to a gate and delay generator (GDG) before going to a bit register. Even with signals from a strip detector and the zero degree detector present, the acquisition system still requires a start command which was controlled by a second CLU, the Master.Live CLU. The second output of the Master CLU is fed into the Master.Live CLU to provide either a start or a veto to the acquisition. The Master.Live CLU allows for a second "good" event condition from the ratedivided zero degree detector events, which does not require an associated proton event. If either a rate-divided zero degree detector event or a proper particle-beam coincidence is seen by the Master. Live CLU, the start of the acquisition is triggered. A "self busy" veto generated when either the acquisition system or the Master.Live CLU is busy may veto the Master. Live output. The Master. Live output also controls the timing sigual gates via a third CLU which requires a signal from the Master.Live CLU and a zero degree detector eveut. The type of event. whether a particle-beam coincidence or a rate-divided zero degree detector event. is distinguished by which bits are set in the data acquisition bit register.


Figure 7.6: Electronics diagram for the beam detectors. Electronics for PPACs 1 and 2 were identical. Gates and time starts were generated by the master electronics shown in Figure 7.8 and are not included explicitly here.


Figure 7.7: Electronics diagram for a single particle telescope. Gates and time starts were generated by the master electronics shown in Figure 7.8 and are not included explicitly here.


Figure 7.8: Diagram of the master electronics which generated event triggers and gates. The boxed number in the coincidence logic units indicates the coincidence level required between the two inputs labeled $A$ and $B$.

## CHAPTER 8

## ANALYSIS AND RESULTS OF $p\left({ }^{20} \mathrm{O}, p\right)^{20} \mathrm{O}^{*}$

### 8.1 Beam Tracking Correction

The inverse kinematics proton scattering angular distribution measurement required considerable off-line analysis to ensure proper particle identification and to accurately determine scattering angles. The PPA('s provided a means for determining the incident beam angle and the interaction position at the target. The position seusitive Si strip detectors, positioned along lines of constant $\theta_{0}$, allow measurement of the scattered proton angle. Beam particle trajectories are not, in general, collinear with the optical beam axis. They are also generally not interacting at the target center. Knowing both the incident beam and the scattered proton trajectories with respect to the optical alignment coordinate system permits a calculation of the final proton scattering angle and the interaction point of the beam with the target. Figure 8.1 shows an arbitrary proton scattering event with the vectors used to recoustruct the beam trajectory and scattering angle.

For determining the beam trajectory, a right-handed Cartesian coordinate system with $\hat{y}$ pointing down (see Figure 8.1) coincident with the optical alignment system proved most useful. An incoming beam particle position is defined at the first PPAC as $\vec{r}_{1}=\left(x_{1}, y_{1}, z_{1}\right)$, and at the second PPAC' by $\overrightarrow{r_{2}}=\left(x_{2}, y_{2}, z_{2}\right)$. The $z_{1}$ and $z_{2}$ components are fixed by the physical placement of the respective PPACs. Knowing these two position measurements, the angle of the beam trajectory with respect to the optical axis is

$$
\begin{equation*}
\theta_{b e a m}=\tan ^{-1}\left[\frac{\sqrt{\left(x_{2}-x_{1}\right)^{2}+\left(y_{2}-y_{1}\right)^{2}}}{\left|z_{2}-z_{1}\right|}\right] . \tag{8.1}
\end{equation*}
$$



Figure 8.1: The trajectory tracking of an arbitrary proton scattering event. The incoming beam particle with trajectory $\vec{r}_{b e a m}$ is detected by the position sensitive PPACs at positions $\vec{r}_{1}$ and $\vec{r}_{2}$ and interacts with a proton in the target at position $\vec{r}_{\text {int }}$. The scattered proton trajectory $\vec{r}_{\text {scat }}$ is at an angle $\theta_{0}$ with respect to the optical axis and an angle $\theta_{\text {scat }}$, the proton laboratory scattering angle, with respect to the incoming beam trajectory. For simplicity, the target is shown perpendicular to the incident beam direction.

The interaction point at the target is then calculated by projecting the beam trajectory onto the target. The target was tilted by approximately $\psi_{t i l t}=50^{\circ}$ with respect to a plane perpendicular to the incident beam direction. For simplicity, Figure 8.1 shows the target perpendicular to the beam direction. The interaction point $\vec{r}_{\text {int }}$ is given by

$$
\begin{align*}
x_{i n t} & =\frac{x_{2}+\frac{\Delta_{x}}{\Delta_{z}} z_{2}}{1-\frac{\Delta_{x}}{\Delta} \frac{1}{\Delta_{2}} \frac{t_{12}}{}}  \tag{8.2}\\
y_{i n t} & =y_{2}+\frac{\Delta_{y}}{\Delta_{z}}\left[z_{2}+\frac{x_{i n t}}{\tan \psi}\right]  \tag{8.3}\\
z_{i n t} & =\frac{x_{i n t}}{\tan \psi} \tag{8.4}
\end{align*}
$$

where $\vec{\Delta}_{x, y, z}=\vec{r}_{2(x, y, z)}-\vec{r}_{1(x, y, z)}$ is the change in beam position between the two PPAC's. A vector describing the beam trajectory $\vec{r}_{b e a m}$ can now be defined as

$$
\begin{equation*}
\vec{r}_{\text {benm }}=\vec{\Delta}_{x, y, z} . \tag{8.5}
\end{equation*}
$$

The position of a scattered proton is measured by its detection at the physical location of one of the Si strips, denoted by a vector $\vec{r}_{\text {strip }}$ from the coordinate system origin to the particular strip. The scattering position is then

$$
\begin{equation*}
\vec{r}_{\text {sent }}=\vec{r}_{s t r i p}-\vec{r}_{i n t} . \tag{8.6}
\end{equation*}
$$

The physical quantity of interest here is $\theta_{\text {scat }}$, which is found by taking the scalar product between the two vectors $\vec{r}_{b e a m}$ and $\vec{r}_{\text {scat }}$ as follows.

$$
\begin{equation*}
\theta_{s c a t}=\cos ^{-1}\left[\frac{\vec{r}_{\text {beam }} \cdot \vec{r}_{\text {scat }}}{\left|\vec{r}_{\text {beam }}\right|\left|\vec{r}_{\text {scat }}\right|}\right] . \tag{8.7}
\end{equation*}
$$

Because the beam profile on target is rather large, the tracking correction is essential in extracting the actual scattering angles. Figure 8.2 is a two dimensional ( $\hat{x}$ and $\hat{y}$ plane) display of the beam projection onto the target. The beam has a FWHM of roughly 2.8 cm , which corresponds to an angular FWHM of $0.9^{\circ}$. An
additional need for the beam tracking arises from the target being tilted. The target tilt of $\psi_{\text {tilt }}=50^{\circ}$ about the $\hat{y}$ axis results in a much larger angular profile of the target being presented to detectors 1,2 , and 3 than to detectors 4,5 , and 6 . This results in a much larger augular correction applied to detectors $1,2,3$ than for the others, because the large beam profile on target expands the angular field of view of each strip in these telescopes. A particularly illuminating way of illustrating this difference is to plot the angular corrections $\Delta \theta=\theta_{\text {scat }}-\theta_{0}$, shown in Figure 8.3. The correction for telescopes 1,2 , and 3 has a FWHM of $4^{\circ}$, while for telescopes 5 . 6 , and 7 , the correction is on the order of the angular size of the individual strips $\left(0.8^{\circ}\right)$, and only a small correction is required.

The gemerated augular correction spectra (Figure 8.3) were used as an off-line gating condition. The gating condition required that $\Delta \theta \leq 1.5^{\circ}$ which limits the angular field of view for each strip in telescopes 1,2 , and 3 to about $3.6^{\circ}$. This degree of correction is used because it represents the minimum correction needed to separate the elastically scattered protons from the inelastically scattered protons. Figure 8.4 shows the effect that the correction has in improving the separation between the elastic and inelastic events. The imposed correction criteria resulted in the loss of about $40 \%$ of the total data from half of the telescopes ( 1,2 , and 3 ) used. The amount of usable data is, however, siguificantly greater than that obtainable had physical collimators been used instead of beam tracking.

A comparison of the beam tracking versus the use of physical collimators can be easily done by imposing software gates on the PPAC spectra which mimic the blocking of the beam by physical collimation. Requiring a beam divergence of $\Delta \theta_{\text {beam }} \leq 0.6^{\circ}$, which is comparable to the angular resolution of the particle telescopes, indicates that roughly $85 \%$ of the total beam events must be discarded. Given that the inteusities of radioactive beams are typically small to begin with, the loss of the majority of the beam is not acceptable. With the use of beam track-


Figure 8.2: Examples of beam tracking spectra show two dimensional position spectra for the beam measured at (a) PPAC 1 and (b) PPAC 2, (c) the projected transverse position, and (d) the $z$ position spectra on target.


Figure 8.3: Angular correction spectra for (a) telescopes 1,2 and 3 and (b) telescopes 5,6 and 7 show the large correction needed for telescopes 1,2 , and 3. The angular resolution for telescopes 5, 6, and 7, however, is on the order of the intrinsic strip width.


Figure 8.4: Proton laboratory kinetic energy versus scattering angle plots from telescopes 1,2 and 3 (a) without and (b) with the beam tracking correction, illustrating the need for a scattering angle correction in order to separate the elastic and inelastic events. The inelastic events are indicated by the dashed boxes.
ing, however, $80 \%$ of the beam is usable while still maintaining acceptable angular resolution.

### 8.2 Proton Angular Distributions

We imposed uumerous software gates in addition to the beam tracking correction to generate the final angular scattering information. As previously mentioned in Chapter 7, a gate placed on the characteristic energy loss of ${ }^{20} \mathrm{O}$ in the zero degree detector ensured that only events associated with an ${ }^{20} \mathrm{O}$ beam particle were accepted. Additionally, $\triangle E-E$ and $T O F-E$ gates were used to filter particle events other than protons seen in the telescopes. The resulting proton scattering events were then sorted into discrete angular bins to generate an angular distribution.

A small portion of the available beam time was spent scattering ${ }^{20} \mathrm{O}$ from a natural carbon target. Since the polypropylene target is composed of $\mathrm{CH}_{2}$, the carbon target data provide a way of subtracting any events from carbon. Due to the very different kinematics of ${ }^{20} \mathrm{O}$ scattering from carbon, though, these events were not in the field of view of the telescope detectors, and were not observed. Hence, no subtraction was needed.

The proton scattering data is most naturally put into angular bins which correspond to the individual strips of the particle telescopes. This method of binning the angular information ensures that all of the bins have an equal solid area, and avoids having to determine efficiencies on a strip by strip basis. In the center of mass, however, this choice of bins results in a measurement of different scattering angles for the elastic and inelastic events. The scattering data is presented in the form of a two dimensional kinematics plot, having energy plotted against the corrected laboratory scattering angle. Figure 8.5 shows the calculated kinematic curves for both the ground state (elastic scattering) and the $2_{1}^{+}$state (inelastic scattering). The same calculation was used to determine the initial placement of the particle telescopes. Figure 8.5 also shows the actual data collected in the experiment. The data shown has been fully corrected for scattering augle. As evidenced by the figure, the data reproduces the expected kinematic curves from the calculations. The data
also shows a clear separation between the ground state and the $2_{1}^{+}$state at 1.67 MeV.

The angular distribution is generated by separately summing the elastic and inelastic counts in a single angular bin (one strip). Two bins per data point were used for the inelastic counts due to lower yields. Representative kinematics plots from a single angular bin, which were used in the angular distribution measurement, are shown in Figure 8.6. A conversion to the center of mass was carried out for each data point consisting of angle and yield information. Since the center of mass angles for a particular strip are different for elastic and inelastic events, this results in a slightly different angular range measured for the two states. The angular distribution in the center of mass is displayed in Figure 8.7. The solid lines through the data points are from a coupled channels analysis of the data, which is discussed in the following section. Error bars denote the width of the angular bin and the uncertainty in the cross section.

### 8.3 Coupled Channels Calculations

Coupled channels calculations (see Chapter 6 for an overview) were performed using the computer code CHUCK [Kunz]. The coupled channels calculations were used to extract the RMS value of the dynamic deformation parameter $\beta_{2}$ of the $2_{1}^{+}$ state in ${ }^{20} \mathrm{O}$ from the experimental elastic and inelastic proton angular distributions. The optical model parameters used in the analysis are from a $35 \mathrm{MeV}\left(p, p^{\prime}\right)$ study by de Swiniarski ct al. [deS74] on the neighboring nucleus ${ }^{20} \mathrm{Ne}$, because the detailed elastic scattering data necessary to perform a credible optical model analysis of ${ }^{20} \mathrm{O}$ is not available. We assume that the parameters for elastic scattering do not vary greatly among neighboring nuclei. The parameters are provided in Table 8.1, and their applicability to ${ }^{20} \mathrm{O}$ is discussed in Chapter 9.


Figure 8.5: Comparison of actual data and calculated kinematics. (a) The collected proton scattering data displayed on the same scale as the calculated kinematics in (b). The dashed box indicates the identified inelastic events. (b) Calculated kinematic curves of energy as a function of scattering angle in the laboratory frame. The solid box shows the field of view of the telescopes. The dashed lines show the punch through energies for a proton penetrating into the next stage of the telescope.


Figure 8.6: Summed proton laboratory kinematics spectra from (a) strip 4 of telescopes 1, 2 and 3 and (b) strip 4 of telescopes 5, 6 and 7. Spectra of this kind were used in summing the cross sections to generate the angular distribution.

As mentioned in Chapter 6, the full optical potential used has both diagonal terms $U_{c c}$ and off diagonal terms $U_{c c}^{\prime}$. $U_{c c}$ has the explicit form

$$
\begin{align*}
U_{c c}= & \frac{2 \mu_{c}}{\hbar^{2}}\left[-V_{R} f\left(x_{R}\right)+i W_{D} \frac{d f\left(x_{I}\right)}{d x_{I}}+V_{\text {Coulomb }}^{\text {Elastic }}\left(r_{c}\right)\right. \\
& \left.+V_{S O} \frac{1}{r_{c}} \frac{d f\left(x_{s O}\right)}{d r_{c}} \vec{l}_{c} \cdot \vec{s}_{c}\right] \tag{8.8}
\end{align*}
$$



Figure 8.7: Angular distributions of protons scattered from the ground state and $2_{1}^{+}$state of ${ }^{20} \mathrm{O}$. The smooth curves are coupled channel calculatious using known optical model parameters from ${ }^{20}$ Ne and the RMS quadrupole deformation $\beta_{2}$ giving the best reproduction of the inelastic cross section.
where $r_{c}$ and $\mu_{c}$ are defined as in Chapter 6 respectively as the relative channel coordinates and the reduced mass of the projectile and target system, and $\vec{l}_{c}$ is the relative angular momentum between the projectile with spin $\vec{s}_{c}$ and the target. The
potential terms contained in $U_{c c}$ are a real Woods-Saxon volume potential $V_{R}$, an imaginary Woods-Saxon surface absorption $W_{D}$, the Coulomb potential $V_{\text {Coulomb }}^{\text {elastic }}$ for elastic scattering, and a spiu-orbit term $V_{s o}$. A volume Woods-Saxon form factor is also used, and is given by

$$
\begin{equation*}
f\left(x_{i}\right)=\frac{1}{\left[1+\epsilon^{x_{i}}\right]} ; \quad x_{i}=\frac{r_{c}-r_{i} A^{\frac{1}{3}}}{a_{i}} \tag{8.9}
\end{equation*}
$$

where A is the target mass number, $V_{i}, r_{i}$, and $a_{i}$ are the potential well depth. radius and diffuseness, respectively, and the subscript $i$ can refer to real (as in $V_{R}$ ), imaginary (as in $W_{D}$ ), or spin-orbit (as in $V_{s o}$ ) potentials as appropriate. The surface Woods-Saxon form factor for the surface absorption $W_{D}$ is taken as the first derivative with respect to $x_{i}$ of the volume form factor.

For the simple vibrational model we use in the present analysis, the off-diagonal (or channel coupling) potential terms $U_{c c^{\prime}}^{\prime}$ are just the first term in an expansion of each potential term in $U_{c c}\left(V_{R}, W_{D}, V_{\text {coulomb }}\right)$ about its equilibrium radius. The spinorbit potential is neglected in the coupling. Assuming a pure quadrupole vibration, so that the $\Psi_{1}^{+}$state is the only state coupled to the elastic channel, gives

$$
\begin{align*}
U_{c c^{\prime}}^{\prime}= & \frac{2 \mu_{c}}{\hbar^{2}}, \beta_{2}\left[V_{R} r_{R} \frac{d f\left(x_{R}^{\prime}\right)}{d x_{R}^{\prime}}-i W_{D} r_{I} \frac{d^{2} f\left(x_{I}^{\prime}\right)}{d x_{l}^{\prime 2}}+r_{c^{\prime}} V_{\text {Coulomb }}^{\text {nelastic }}\left(r_{c^{\prime}}\right)\right] \\
& \times \sum_{m} D_{m}^{2}\left(\epsilon_{t}\right) Y_{2 m}\left(\hat{r}_{c^{\prime}}\right), \tag{8.10}
\end{align*}
$$

where $\beta_{2}$ is the RMS quadrupole deformation of the target in the $2_{1}^{+}$state inelastic channel. The $D_{m}^{2}\left(\epsilon_{t}\right)$ performs a rotation through the Euler angles of $Y_{2 m}$, and the $\epsilon_{t}$ represent the Euler angles connecting the intrinsic reference frame of the target to the channel coordinates.

Besides defining the optical model parameter values, the computer code CHUCK also allows definition of multiple deformation parameters for every channel included in the coupling. For our simple vibrational model in the present analysis, only the RMS quadrupole deformation $\beta_{2}$ of the $2_{1}^{+}$state was used. For simplicity, the

Table 8.1: Optical model parameters for ${ }^{20} \mathrm{Ne}$ and ${ }^{20} \mathrm{O}$

| $V_{R}$ | $r_{R}$ | $a_{R}$ | $W_{L}$ | $r_{I}$ | $a_{I}$ | $V_{S O}$ | $r_{S O}$ | $a_{S O}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $[\mathrm{MeV}]$ | $[\mathrm{fm}]$ | $[\mathrm{fm}]$ | $[\mathrm{MeV}]$ | $[\mathrm{fm}]$ | $[\mathrm{fm}]$ | $[\mathrm{MeV}]$ | $[\mathrm{fm}]$ | $[\mathrm{fm}]$ |
| 44.3 | 1.10 | 0.73 | 6.96 | 1.38 | 0.60 | 7.80 | 1.03 | 0.74 |

calculation was done for standard 30 MeV proton scattering kinematics, where the protons were treated as projectiles and ${ }^{20} \mathrm{O}$ as stationary target nuclei. This makes no difference to the results, since the calculation is carried out in the center of mass frame.

A test of our coupled channels calculation method was carried out to reproduce the results of de Swiniarski et al. [deSit4] for ${ }^{20} \mathrm{Ne}$ prior to our calculations for ${ }^{20} \mathrm{O}$. The results of de Swiniarski ct al. were performed using a rotational model for ${ }^{20}$ Ne with the computer code E(IS. Using the identical uptical model parameters (see Table 8.1), we were able to reproduce the results of de Swiniarski ct al. by using a vibrational model in our coupled channels calculations. Although a rotational model is certainly more applicable to ${ }^{20} \mathrm{Ne}$, this test indicates that our coupled channels calculations are not sensitive to the choice of rotational or vibrational models. A similar insensitivity was reported by Grabmayr et al. [CiraS0] in their analysis. which also used ('HU(KK, of inelastic nucleon scattering for the $2_{1}^{+}$state in ${ }^{18} \mathrm{O}$.

Our calculations for ${ }^{20} \mathrm{O}$ were performed in an identical manner to the test calculations. We adjusted the magnitude of $\beta_{2}$ in successive, but otherwise identical, calculations to obtain the best reproduction of the inelastic scattering cross section. The integration was performed with 20 partial waves out to a distance of 10 fm in steps of 0.1 fm . The best reproduction of the $2_{1}^{+}$state angular distribution is with $\beta_{2}=0.50 \pm 0.04$, which is very close to the $\beta_{2}$ value of 0.47 in ${ }^{20}$ Ne reported by de Swiniarski et al. [deS74]. The calculation results and the experimentally measured angular distribution are shown in Figure S.ī.

## 8.4 $M_{n} / M_{p}$ Determination

Calculating $M_{n} / M_{p}$, the ratio of the ueutron to proton multipole matrix elements, for the $0_{g s}^{+} \rightarrow 2_{1}^{+}$transition in ${ }^{20} \mathrm{O}$ follows the method put forth in References [Mad75, Ber79, Ber81, Ber8:3]. The two experimental values required in the calculation are the RMS quadrupole deformation parameter $\beta_{2}$ from ( $p, p^{\prime}$ ), determined from the proton scattering angular distribution, and the reduced transition probability $B(E 2)$ from an electromagnetic probe. We use the adopted value [Ram87] of $B(E 2)=1.80 \pm 0.07 \mathrm{~W} . u$. from lifetime measurements of the $\gamma$-ray decays from the $2_{1}^{+}$state [Ber75, HeriT, Rut80, Til95]. Following the prescription of Bernstein et al. [Ber83], $M_{n} / M_{p}$ is determined from these two experimental quantities by the relation

$$
\begin{equation*}
\frac{M_{n}}{M_{p}}=\frac{b_{p}}{b_{n}}\left[\frac{\delta_{\left(p, p^{\prime}\right)}}{\delta_{e m}}\left(1+\frac{b_{n}}{b_{p}} \frac{N}{Z}\right)-1\right] . \tag{8.11}
\end{equation*}
$$

The terms $b_{n}$ and $b_{p}$ denote the sensitivities of the particular experimental probe to neutron and proton distributions respectively. As discussed previously (see Chapter 6), proton scattering in the energy range of the present experiment ( $10-50 \mathrm{MeV}$ ) is roughly three times more sensitive to the neutron distribution as to the protons, as given in Table 6.1. This gives $b_{n} / b_{p} \approx 3$ in the present work. The terms $\delta_{\left(p, p^{\prime}\right)}$ and $\delta_{\text {em }}$ are respective deformation lengths for proton scattering and electromagnetic probes, also discussed in Chapter 6. We assume a uniformly spherical neutron distribution to relate $\delta_{\left(p, p^{\prime}\right)}$ to the RMS quadrupole deformation value $\beta_{2}\left(p, p^{\prime}\right)$ by

$$
\begin{equation*}
\delta_{\left(p, p^{\prime}\right)}=\beta_{2}\left(p, p^{\prime}\right) r_{0} A^{1 / 3} \tag{8.12}
\end{equation*}
$$

where $r_{0}=1.1 \mathrm{fm}$ is used for consistency with the real part of the optical potential ([deS74] and Table 8.1). Likewise, since this is a simple collective model, we assume a uniformly spherical proton distribution, and relate $\delta_{e m}$ to $B(E 2)$ by the relation

$$
\begin{equation*}
\delta_{e m}=\frac{r_{C} A^{1 / 3}}{5 Z}\left[B\left(E^{2}\right) 20 \pi\right]^{1 / 2} \tag{8.13}
\end{equation*}
$$

where $r_{C}=1.2 \mathrm{fm}$ (the Coulomb radius) is used. This gives, finally,

$$
\begin{equation*}
\frac{M_{n}}{M_{p}}=\frac{1}{3}\left[5 Z\left(\frac{1.1}{1.2}\right) \frac{\beta_{2}\left(p, p^{\prime}\right)}{\sqrt{20 \pi B(E 2)}}\left(1+\frac{3 N}{Z}\right)-1\right] \tag{8.14}
\end{equation*}
$$

Substituting the values of $\beta_{2}\left(p, p^{\prime}\right)$ and $B\left(E^{2}\right)$ gives $M_{n} / M_{p}=2.8 \pm 0.4$.
Bernstein et al. have pointed out that this approach of obtaining $M_{n} / M_{p}$ is somewhat model dependent [Ber81, Ber8:3]. However, the difference in most nuclear models reveals itself primarily in the transition densities $\rho_{f_{i}( }(r)$ for small values of $r$, and proton scattering at 30 MeV is not very sensitive to the nuclear interior [Ber8:3]. Carr et al. pointed out that mucleon scattering of at least 60 MeV was needed to effectively probe the muclear interior [Car85b], due to the attenuation of the uucleon wave. Berustein ct al. also point out that the $r^{-1}$ dependence of the multipole operator $O_{\lambda \mu}^{F}$, which is decomposed into neutron and proton operators in Equation 6.3, reduces the sensitivity of the probe to the nuclear interior [Ber8:3]. Furthermore, our measured proton angular distribution is primarily at forward angles ( $20^{\circ}$ to $45^{\circ}$ ), where model dependence is less important. The effects from model dependence have been shown to be small for low energy $\alpha$ scattering [Ber69, Mad77] and proton scattering [Wami9].

## CHAPTER 9

## DISCUSSION OF ${ }^{20} \mathrm{O}$ RESULTS

The present work finds that $M_{n} / M_{p}=2.8 \pm 0.4$ for the $0_{g s}^{+} \rightarrow 2_{1}^{+}$transition in ${ }^{20} \mathrm{O}$ is significantly higher than the $N / Z=1.5$ value expected for a purely isoscalar excitation. The high $M_{n} / M_{p}$ value shows that the $2_{1}^{+}$excitation is dominated by neutron oscillations from the four valence neutrons in the $2 s 1 d$ shell that lie outside the relatively inert core of completely filled $Z=S$ proton and $N=S$ neutron shells. This result is consistent with our expectations for a singly closed shell nucleus. Hence, it appears that the underlying shell structure of this nucleus plays a large role in the $0_{g s}^{+} \rightarrow \nu_{1}^{+}$transition, and that the simple picture of a collective, isoscalar excitation is not a sufficient description.

The systematic behavior of $\left(M_{n} / M_{p}\right) /(N / Z)$ for $0_{g s}^{+} \rightarrow 2_{1}^{+}$transitions in the $12 \leq A \leq 26$ mass region can be seen in Figure 9.1. Results from proton scattering [DeL83, GraS0, Ril:diss, deS74, Has83, Zwi8:3], neutron scattering [Gra80, Ols89, Ols90], and electromagnetic [Ram87] measurements are included where available. The proton and neutron scattering results are compiled from sources where a coupled channels analysis similar to that used in the present work was used to extract RMS quadrupole deformation parameters. In the case of ${ }^{18} \mathrm{Ne}$ [Ril:diss], the value was obtained from the same experimental apparatus and analysis method as the present work. All of the reported $M_{n} / M_{p}$ values were derived from Equation 6.7 with the appropriate probe sensitivities from Table 6.1. The dashed lines in Figure 9.1 are shell model predictions from Brown et al. [Bro82], and the solid line represents $M_{n} / M_{p}=N / Z$, the expected value for equal neutron and proton oscillation amplitudes.


Figure 9.1: Ratios of $\left(M_{n} / M_{p}\right) /(N / Z)$ for $0_{g s}^{+} \rightarrow 2_{1}^{+}$transitions in the $12 \leq A \leq$ 26 mass region. The values are extracted from comparisons of proton scattering [DeL83, Gra80, Ril:diss, deS74, Has833, Zwi83], neutron scattering [Gra80, Ols89, Ols90], and electromagnetic [Ram87] measurements. The dashed lines are shell model predictions from [Bro82].

Although the shell structure appears to have a large effect on the $2_{1}^{+}$excitation in ${ }^{20} \mathrm{O}$, Figure 9.1 suggests that the simple shell predictions of Reference [Bro82] do not provide an adequate description of the entire mass region. While we expect that singly closed shell nuclei would be better described with a shell model than would
open shell nuclei, the results of Brown et al. [Bro82] do not adequately explain the data for both categories of nuclei. The shell model predictions are very close to the experimental $M_{n} / M_{p}$ values in the open shell nucleus ${ }^{26} \mathrm{Mg}$ and the singly closed shell nucleus ${ }^{18} \mathrm{Ne}$. However, the predictions are substantially different from the experimental values for the open shell uucleus ${ }^{20} \mathrm{Ne}$ and the singly closed shell nucleus ${ }^{18} \mathrm{O}$. Brown et al. have pointed out that the most likely cause for errors in their calculations (described in Chapter 6) stems from their use of a harmonic oscillator radial dependence in the single-particle wave functions [Bro82]. They also point to questionable experimental results, and stress the importance of confirming previous measurements. Although there is good agreement between the combined proton scattering and electromagnetic method ( $p-E M$ method) and the combined proton and neutron scattering method ( $p-n$ method), we see discrepancies in the reported $M_{n} / M_{p}$ values when compared to results derived from pion scattering.

Figure 9.2, very similar to Figure 9.1. is a plot of $\left(M_{n} / M_{p}\right) /(N / Z)$ taken from pion scattering results compiled by Peterson [Pet93]. The figure compares $M_{n} / M_{p}$ values obtained from scattering both $\pi^{+}$and $\pi^{-}$pions from fixed targets with the values obtained from $p-E M$ and $p-n$ methods. Pion scattering, as mentioned in Chapter 1 , is au indepeudeut method of measuring $M_{n} / M_{p}$.

The pion scattering $M_{n} / M_{p}$ values given bere agree qualitatively with the simple shell model picture, where deviations from the simple collective $M_{n} / M_{p}=N / Z$ value are expected for the singly closed shell nuclei ${ }^{14} \mathrm{C}$ and ${ }^{18} \mathrm{O}$. Likewise, the $M_{n} / M_{p}$ values fall close to the $N / Z$ expectation for the open shell nuclei ${ }^{20,22} \mathrm{Ne}$ and ${ }^{24,26} \mathrm{Mg}$. However, the quantitative shell model predictions of Brown et al. [Bro82], do not agree with the $M_{n} / M_{p}$ results derived from pion scattering any more than they do for the results derived from the $p$-EM or $p-n$ methods. In fact, in the three nuclei where a direct comparison is possible $\left({ }^{18} \mathrm{O},{ }^{20} \mathrm{Ne}\right.$, and $\left.{ }^{26} \mathrm{Mg}\right)$, the shell model predictions fall in between the $M_{n} / M_{p}$ values derived from piou scattering and nucleon scattering.


Figure 9.2: $\left(M_{n} / M_{p}\right) /(N / Z)$ ratios for the $0_{g s}^{+} \rightarrow 2_{1}^{+}$transitions in the $12 \leq A \leq 26$ mass region. The results from pion scattering are compared to those from nucleon scattering and electromaguetic results. The dashed lines are shell model predictions from [Bro82].

The discrepancy between the results obtained by pion scattering and those obtained from uucleon scattering raises some concern. When viewed as a whole in the $12 \leq A \leq 26$ mass region, the results obtained by pion scattering seem to best match our expectations from simple shell model and collective model arguments.

Even though $M_{n} / M_{p}$ results derived from the pion scattering and nucleon scattering agree well for $A \geq 40$ [Ber81], many authors have pointed out the discrepancy in the $12 \leq A \leq 26$ mass region [Ber81, Bro82, Ber83, Car85b, Pet93]. Unfortunately, a satisfactory answer to this discrepancy has not yet been given. Both pion scattering near the $\Delta$ resonance and low energy nucleon scattering are strongly absorbed at the uuclear surface. Although low energy nucleons can penetrate more deeply into the nuclear interior than pions at the $\Delta$ resonance, effects from this are expected to be very slight for incident nucleons below 60 MeV [Car85b].

It is also quite possible that we have not entirely removed the probe sensitivity from the present analysis. As discussed in Chapter 6, the probe sensitivity factors we have used from Reference [Ber81] are based on the assumption that the isoscalar strength of the optical potential is twice that of the isovector strength. This is a crude assumption: however, it has been used with much success in $A \geq 40$ nuclei, where very little discrepancy between $M_{n} / M_{p}$ results derived from pion scattering and nucleon scattering is seen [Bers1].

Additionally, a systematic bias may exist in this mass region in the deformed optical model potential method most often used to extract $\beta_{2}$ from cross-section data. The piou scattering $M_{n} / M_{p}$ value for ${ }^{18} \mathrm{O}$. for example, is significantly higher than that obtained from nucleon scattering. Indeed, with the exception of ${ }^{12} \mathrm{C}$ where the values nearly agree, a higher $M_{n} / M_{p}$ value is obtained from pion scattering in every nucleus ( ${ }^{18} \mathrm{O},{ }^{20,22} \mathrm{Ne}$, and ${ }^{24,26} \mathrm{Mg}$ ) where a direct comparison is possible. The fact that the results derived from pion scattering are always larger would seem to indicate a systematic difference in the two analyses. Since there are so few data points with which to make a direct comparison, however, we can not say with certainty whether or not a systematic analysis difference exists. All of the pion scattering results have been taken from a consistent analysis [Pet93] in which a folding calculation with a distorted wave impulse approximation (see Reference
[Sat83]) was used. A similar, consistent analysis of the nucleon scattering data using a folding model analysis would be quite helpful in detecting a systematic analysis difference.

Since neither $M_{n} / M_{p}$ values from pion scattering nor shell model calculations exist for the $0_{g s}^{+} \rightarrow 2_{1}^{+}$trausition in ${ }^{20} \mathrm{O}$, it is difficult to evaluate the extent of any systematic analysis error in the present work. We can, however, examine possible sources of error in our experiment and our subsequent analysis. The scattered proton angular distribution we observe is limited by the angular resolution of the detectors $\left(0.8^{\circ}-3.6^{\circ}\right)$ and by our ability to separate the elastically scattered from the inelastically scattered protons. Each of these factors contributes ouly slightly to the total error of about $20 \%$ in our determination of $M_{n} / M_{p}$. This is because the extraction of the RMS deformation parameter $\beta_{2}$ from our coupled channels calculations is primarily dependent upon the total yield of inelastic events relative to elastic events. Roughly $80 \%$ of the total error in our $M_{i n} / M_{p}$ determination comes from uncertainty in the $\gamma$-ray lifetime measurement of $\mathrm{B}(\mathrm{E} 2)$. The angular distribution over a range of $15^{\circ}-45^{\circ}$ in the center of mass, does, however allows us to check the reasonableness of our coupled channels calculations.

As stated in Chapter 8 , we used the existing optical model parameters from the nucleus ${ }^{20} \mathrm{Ne}$ [deS74] in our coupled channels analysis of the proton angular distribution from ${ }^{20} \mathrm{O}$. The elastic scattering data for ${ }^{20} \mathrm{O}$ is not sufficient to extract reliable optical model parameters, and we chose the parameters of ${ }^{20} \mathrm{Ne}$ as the best alternative for our aualysis. The best indication that those optical model parameters are applicable to ${ }^{20} \mathrm{O}$ is that they reproduce well the observed angular distribution for the elastically scattered protons (see Figure 8.7). Furthermore, we have used a phenomenological optical model which is based on the assumption that the optical model parameters vary slowly and smoothly throughout the Periodic Table [Sat83],
so that we don't expect large differences in the optical model parameters for elastic scattering between neighboring nuclei.

Even with the possibility of a systematic discrepancy between the pion scattering analysis and the proton scattering analysis, we can conclude that the $0_{g s}^{+} \rightarrow 2_{1}^{+}$ transition in ${ }^{20} \mathrm{O}$ is dominated by valence neutron contributions consistent with $Z=8$ proton shell closure effects. The $p-E M$ method of the present experiment has been used extensively and successfully in the $A \geq 40$ mass region where $M_{n} / M_{p}$ values agree with those derived from pion scattering [Ber81]. If, as it seems from the available data, the pion scattering analysis yields a systematically larger $M_{n} / M_{p}$ value than the proton scattering analysis, our result of $M_{n} / M_{p}=2.8 \pm 0.4$ from the $p-E M$ method is clearly an indication of a dominant neutron contribution to the $0_{g s}^{+} \rightarrow 2_{1}^{+}$transition. We can not claim, however, to have made a precise measurement of $M_{n} / M_{p}$, as evidenced by the large error bars in Figure 9.2. Nor can we yet claim that $M_{n} / M_{p}$ is getting larger as we move farther from stability towards the neutron drip line. Such a conclusion would require higher quality data.

In order to understand the systematic behavior of $M_{n} / M_{p}$ for the $0_{g s}^{+} \rightarrow 2_{1}^{+}$ transitions in nuclei approaching the drip line, additional experimental information is required. The singly closed shell uuclei offer the best possibility for confirmation of a large valence nucleon contribution observed in ${ }^{18,20} \mathrm{O}$ and ${ }^{18} \mathrm{Ne}$. A study similar to the present one on the uuclei ${ }^{22,24} \mathrm{O}$ would allow systematic comparisons of the present data with other $Z=8$ closed proton shell nuclei. The intensities for producing these secondary beams have been experimentally determined [Fa:diss] at $\approx$ 300 pps and $\approx 1 \mathrm{pps}$ respectively, at the present NSCL facility, well below the desired intensities of a few thousand pps for successful inverse kinematics proton scattering experiments. However, the planned coupled cyclotron upgrade at the NSCL should increase the available secondary beam intensities by an estimated factor of 1,000 or
more, making studies for other singly closed shell nuclei in the $12 \leq A \leq 26$ mass region quite feasible.

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## CHAPTER 10

## CONCLUSION

Two different proton scattering experiments have been performed to study low energy vibrational excitations. Although the two experiments both involve proton scattering, the techniques used and the information obtained is very different for the two.

A $3{ }_{2}^{-}$state at 2423.4(4) keV was rigorously identified in the even-even Pt isotope ${ }^{196} \mathrm{Pt}$. The state was populated via a ${ }^{196} \mathrm{Pt}\left(p, p^{\prime} \gamma\right)$ proton- $\gamma$ coincidence experiment with a 12.7 MeV proton beam at Florida State University. The spin assignment is based on previous aualysis of ( $p, p^{\prime}$ ) angular distributions [Cot88b], which limit the allowed spins to $J^{\pi}=3^{-}, 4^{+}$, and previous observation of the 2423.4 keV state in ( $n, n^{\prime} \gamma$ ) [DiP9:3], which excludes the $J^{\pi}=4^{+}$assignment. We therefore assign a spin of $J^{\pi}=3^{-}$to the 2423.4 keV state, confirming a previous assignment by Ponomarev et al. [Pon92]. The observation of a $\gamma$-ray decay of $976.7(3) \mathrm{keV}$ to the lowest-lying $3^{-}$state at $1447.0(2)$ keV confirms the placement of the level.

Observation of a higher-lying $3^{-}$state considerably strengthens the argument made in Reference [Cot88b] for fragmentation of the low energy octupole state in ${ }^{196} \mathrm{Pt}$. Three additional states, not observed in the present experiment, have nearly identical proton angular distributions, and it is likely that they, too, have $J^{\pi}=3^{-}$ and are additional fragments of the low energy octupole state [Cot88b]. Ponomarev et al. has already assigned $J^{\pi}=3^{-}$to two of these states [Pon92]. Furthermore, a similar analysis of protou angular distributions from ${ }^{194,198} \mathrm{Pt}$ [Cot88b] shows a number of states in the two isotopes with nearly identical distributions to the 2423.4 keV state in ${ }^{196} \mathrm{Pt}$. It is likely that the octupole state is fragmented in these nuclei,
also. Transition strengths on the order of 6-9 W.u. indicate a moderate degree of collectivity in these states, identifying them as collective octupole vibrations. ${ }^{196} \mathrm{Pt}$ is one of the best known examples of an $\mathrm{O}(6)$ symmetry nucleus [Ciz78] in the IBM formalism, in which the apparent octupole fragmentation is understood. Similar ${ }^{194,198} \operatorname{Pt}\left(p, p^{\prime} \gamma\right)$ experiments are needed to identify octupole fragmentation in the other even-even Pt isotopes.

Proton scattering is used for the first time to excite the $2_{1}^{+}$state in the neutron rich ${ }^{20} \mathrm{O}$ uucleus. In this experiment, however, the proton scattering is performed in inverse kinematics from the radioactive beam ${ }^{20} \mathrm{O}$ at the NSCL at MSU. The RMS quadrupole deformation parameter, $\beta_{2}$, of the $2_{1}^{+}$excitation at 1.674 MeV is extracted with a coupled channels analysis of the measured elastic and inelastic scattered proton angular distributions. Used in combination with the reduced transition probability $\mathrm{B}(\mathrm{E} 2)$ taken from the known $\gamma$-ray lifetime [Ram87], the ratio of neutron to proton multipole matrix elements is determined as $M_{n} / M_{p}=2.8 \pm 0.4$.

The result is significantly larger than the isoscalar estimate of $M_{n} / M_{p}=N / Z=$ 1.5 for ${ }^{20} \mathrm{O}$, indicating a substantially larger neutron contribution relative to the proton contribution in the $0_{g s}^{+} \rightarrow 2_{1}^{+}$transition. An enhanced neutron contribution is consistent with a simple shell model picture where the four valence neutrons outside of a closed shell core ( $N=S$ and $Z=S$ ) dominate the excitation. A similar enhanced ueutron contribution was observed in ${ }^{18} \mathrm{O}$ [Pet93], which has the same $Z=8$ closed proton shell. Additional measurements of other singly closed shell nuclei in the $12 \leq A \leq 26$ mass region would allow a systematic determination of the behavior of $M_{n} / M_{p}$ for $0_{g s}^{+} \rightarrow 2_{1}^{+}$transitions near the neutron and proton drip lines. Systematic studies would also lend insight into the discrepancies observed between pion scattering and nucleon scattering results in this mass region.

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## BIOGRAPHICAL SKETCH

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